



CP Violation for the Heavens and the Earth

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Oct. 12 @ CFPP 2021, NCTS/NTU, Taipei



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CP Violation for the Heavens and the Earth



- 0. Intro: Our Life & Times
- I. **General 2HDM**
- II. Electroweak Baryogenesis
- III. Under the Heavens on Earth: **e EDM**
- IV. Phenomenological Consequences
- V. Summary



0. Intro: Our Life & Times

$h(125)$ ✓ *New Physics* ✗

Where is SUSY?

Out-of-the-box Searches (XLPs) & EFTs ⇐ Where is Everybody?

Assuming only the SM states below some Λ :



We can then Taylor expand (SM fields and derivative over Λ):

$$\mathcal{L}_{\text{eff}} = \frac{\Lambda^4}{g_*^2} \mathcal{L} \left(\frac{D_\mu}{\Lambda}, \frac{g_H H}{\Lambda}, \frac{g_{f_{L,R}} f_{L,R}}{\Lambda^{3/2}}, \frac{g_{F_{\mu\nu}}}{\Lambda^2} \right) \simeq \mathcal{L}_4 + \mathcal{L}_6 + \dots$$

(assuming lepton & baryon number)

dimension-4 terms:

The SM

dimension-6 terms:

Leading deviations from the SM



0. Intro: Our Life & Times

$h(125)$ ✓ *New Physics* ✗

Where is SUSY?

Out-of-the-box Searches (XLPs) & EFTs ← Where is Everybody?

unconventional-Conventional

Assuming only the SM states below some Λ :



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(assuming lepton & baryon number)

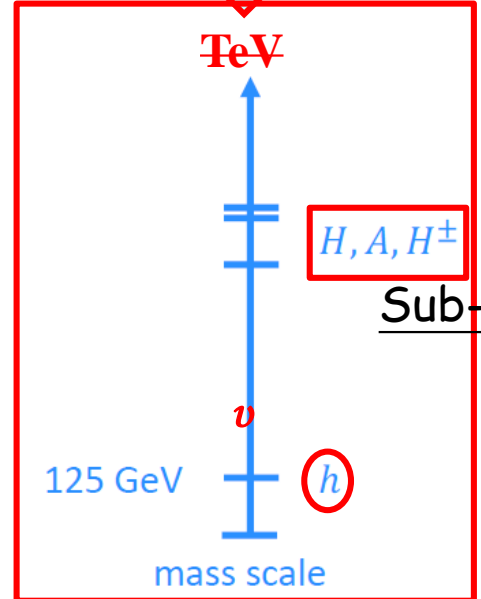
dimension-4 terms

The SM (G2HDM) General 2HDM

dimension-6 terms: Leading deviations from the SM

Road Not Taken

Extra Higgs Doublet w/ Extra Yukawa Couplings & Extra Quartic Couplings





0. Intro: Our *Life* & Times

$h(125)$ ✓ *New Physics* ✗



However,

Where is SUSY?

Beyond CKM CPV

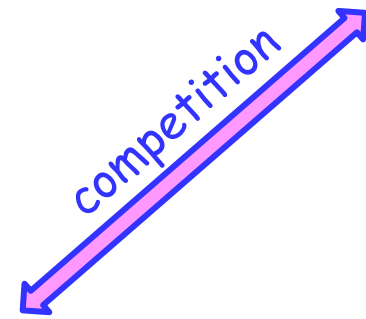
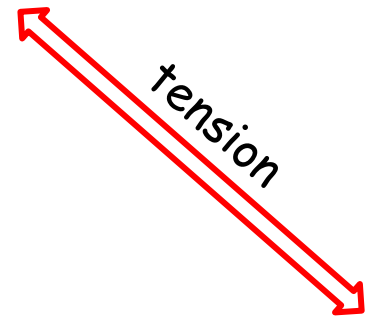
EW BaryoGenesis (EWBG)

- more testable -



LHC

- *No New Physics* -



eEDM: ACME14 → ACME18

- L.E. Precision Frontier -

$$|d_e| < 1.1 \times 10^{-29} e \text{ cm}$$

Evolution of the Universe

15 billion years

life on earth



molecules form

1 billion years

heavy elements formed in stars

The Universe began with a "Big Bang" ~ 14 billion years ago

1 million years

stars and galaxies exist,

300,000 years

microwave background radiation fills universe

atoms form

3 minutes

helium nuclei formed

1 second

neutrons and protons formed

10s quark "soup" matter dominates

10^{15} deg

10^{10} deg

10^9 deg

6000°

4000°

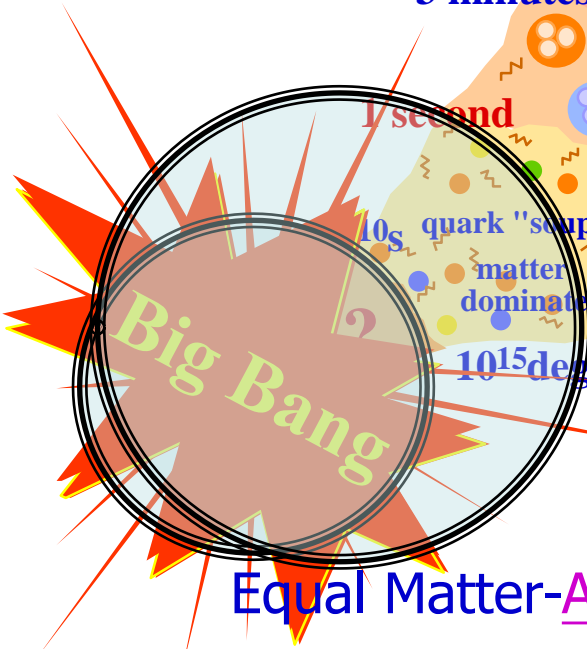
-255°

270°

Matter Only !?



Equal Matter-Antimatter





Soaring to the Starry Heavens



Enough **CPV**
for **BAU**?



天

CPV → **BAU**

地



ACME experiment: Current frontline, Probe **CPV** via *eEDM*, put check on **Baryogenesis**.



I. General 2HDM

Two Higgs Doublet Model

Whither 1st Order Phase Trans. / *Sufficient CPV?*

SM: Weak Int. *too Weak* / Jarlskog Invariant *way too small!*
All 3 gens. \Rightarrow Mass and CKM suppressed

2HDM: $\mathcal{O}(1)$ Higgs Quartics OK / CPV in $V(\Phi_1, \Phi_2)$ problematic w/ d_n
Wise to keep $V(\Phi_1, \Phi_2)$ CP Conserving

Comment: Known CPV in CKM, i.e. Yukawa's. Extra Yukawa's?

Killed by Z_2 (Glashow-Weinberg 1977) for Flavor Conservation.

General

2HDM: $\mathcal{O}(1)$ Higgs Quartics OK / Extra Yukawa's w/o Z_2

$\mathcal{O}(1)$
 ρ_{tt} the driver; ρ_{tc} the backup

N.B. Data-driven ρ_{ij} : $t \rightarrow ch$; $h \rightarrow \mu\tau \dots$

Fuyuto, WSH, Senaha, Phys. Lett. B 776 (2018) 402



Extra Yukawa Couplings

General 2HDM w/o Z_2

General Yukawa interaction for up-type quarks

$$-\mathcal{L}_Y = \bar{q}_{iL} (Y_{1ij}^u \tilde{\Phi}_1 + Y_{2ij}^u \tilde{\Phi}_2) u_{jR} + \text{h.c.}$$

$v_1 = v c_\beta \quad v_2 = v s_\beta$

$Y^{\text{SM}} = Y_1 c_\beta + Y_2 s_\beta$

$m_f = y_f v / \sqrt{2}$

$V_L^{u\dagger} Y^{\text{SM}} V_R^u = \text{diag}(y_u, y_c, y_t) \equiv Y_D \quad \text{diagonal}$

$\rho = V_L^{u\dagger} (-Y_1 s_\beta + Y_2 c_\beta) V_R^u$

FCNH (flavor changing neutral H)

Neutral up-type Yukawa interaction

$$-\mathcal{L}_Y = \bar{u}_{iL} \left[\frac{y_i \delta_{ij}}{\sqrt{2}} s_{\beta-\alpha} + \frac{\rho_{ij}}{\sqrt{2}} c_{\beta-\alpha} \right] u_{jR} h$$

$$+ \bar{u}_{iL} \left[\frac{y_i \delta_{ij}}{\sqrt{2}} c_{\beta-\alpha} - \frac{\rho_{ij}}{\sqrt{2}} s_{\beta-\alpha} \right] u_{jR} H$$

$$- \frac{i}{\sqrt{2}} \bar{u}_{iL} \rho_{ij} u_{jR} A + \text{h.c.},$$

$c_{\beta-\alpha} \rightarrow 0$

alignment limit!

→ diag.

FCNH ρ_{ij}

$|\rho_{ij}| e^{i\phi_{ij}}$



Extra Higgs Quartic Couplings

Sub-TeV Spectrum & 1stEWPT

SM

$$V(\Phi) \sim -\mu^2|\Phi|^2 + \lambda|\Phi|^4$$

$$v^2 \sim \mu^2/\lambda$$

G2HDM

$$V(\Phi, \Phi') = \mu_{11}^2|\Phi|^2 + \mu_{22}^2|\Phi'|^2 - (\mu_{12}^2\Phi^\dagger\Phi' + \text{h.c.}) + \frac{\eta_1}{2}|\Phi|^4 + \frac{\eta_2}{2}|\Phi'|^4 + \eta_3|\Phi|^2|\Phi'|^2 + \eta_4|\Phi^\dagger\Phi'|^2 + \left\{ \frac{\eta_5}{2}(\Phi^\dagger\Phi')^2 + [\eta_6|\Phi|^2 + \eta_7|\Phi'|^2]\Phi^\dagger\Phi' + \text{h.c.} \right\}$$

WSH&Kikuchi, EPL'18

$$\mu_{12}^2 = \frac{1}{2}\eta_6 v^2$$

"min. cond."

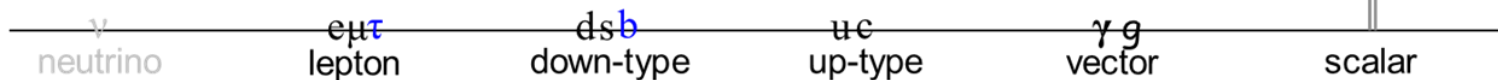
η_6 : sole param. for h-H mixing (c_γ)

Dim'less params. $\mathcal{O}(1)$ ("Common" Naturalness):

$$\eta_i \text{ with } i = 1-7; \mu_{22}^2/v^2$$



Search Zone



a Sakharov condition: 1stEWPT

N.B. $\mathcal{O}(1)$ η_i 's needed for 1st order Phase Trans., prerequisite for ElectroWeak BaryoGenesis.



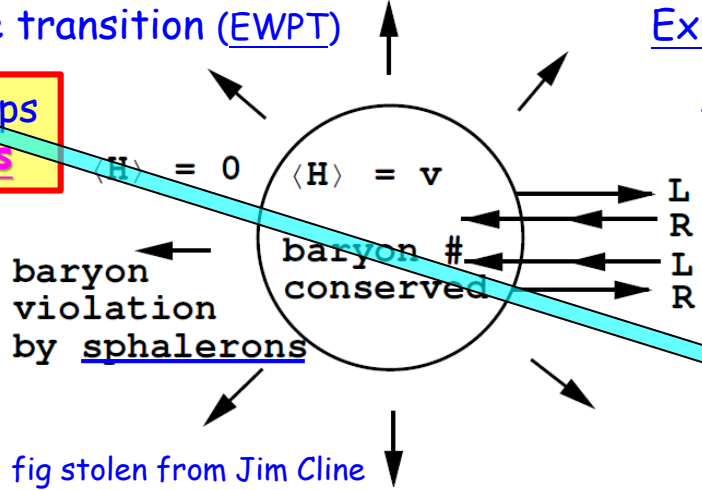
II. EWBG

strongly 1st order EW phase transition (EWPT) ↑

Expanding Bubble of Broken Phase ↑

Extra Higgs Thermal Loops
w/ $\mathcal{O}(1)$ Higgs Quartics

2HDM OK



To avoid n_B washout:
Hubble const.

$$\Gamma_B^{(br)}(T_C) < H(T_C)$$

n_B changing rate (br)

$$v_C/T_C > \zeta_{sph}(T_C) \sim \mathcal{O}(1)$$

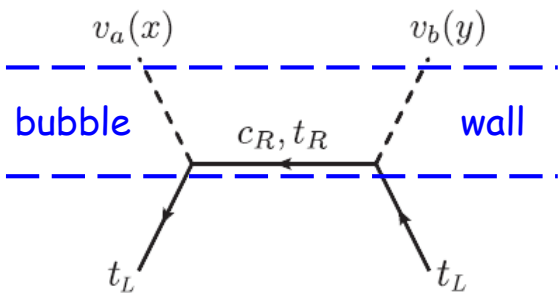
$$v_{ev} @ T_C = \sqrt{v_1^2(T_C) + v_2^2(T_C)}$$

Baryon Asymm. of Universe (BAU)

n_B/s

$$Y_B = \frac{-3\Gamma_B^{(sym)}}{2D_q\lambda_+s} \int_{-\infty}^0 dz' n_L(z') e^{-\lambda_- z'}$$

Planck 2014
 $Y_B^{obs} = 8.59 \times 10^{-11}$



$$\Gamma_B^{(sym)} = 120\alpha_W^5 T$$

$$D_q \simeq 8.9/T$$

s

$$\lambda_{\pm} \simeq v_w$$

$$n_L$$

z'

n_B changing rate (sym)

quark diffusion const

entropy density

bubble wall velocity

l.h. fermion density (l.h. top density)

coord. oppo. bubble exp. dir.



CPV Top interactions

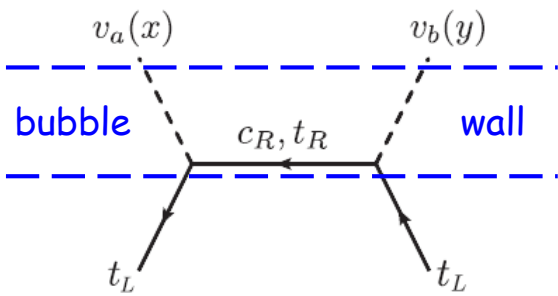
CPV source term

$$S_{i_L j_R}(Z) = N_C F \text{Im} \left[(Y_1)_{ij} (Y_2)_{ij}^* \right] v^2(Z) \partial_{t_Z} \beta(Z)$$

- $Z = (t_Z, \mathbf{Z})$ position in heat bath (Very Early Univ.)
- $N_C = 3$ # of color
- F function of complex energies for i_L, j_R
- $\partial_{t_Z} \beta(Z)$ physical variation ($\Delta\beta = 0.015$)

Baryon Asymm. of Universe (BAU) n_B/s

$$Y_B = \frac{-3\Gamma_B^{(\text{sym})}}{2D_q \lambda_{+s}} \int_{-\infty}^0 dz' n_L(z') e^{-\lambda_- z'}$$



BAU \leftarrow CPV Top interactions at Bubble Wall

left-handed Top density



n_L

skip detail (Transport)

z'

coord. oppo. bubble exp. dir.



CPV Top interactions

CPV source term

$$S_{iLjR}(Z) = N_C F \text{Im} [(Y_1)_{ij} (Y_2)_{ij}^*] v^2(Z) \partial_{t_Z} \beta(Z)$$

$$\text{Im} [(Y_1)_{ij} (Y_2)_{ij}^*] = \text{Im} [(V_L^u Y_D V_R^{u\dagger})_{ij} (V_L^u \rho V_R^{u\dagger})_{ij}^*]$$

To understand the plot to follow, suppose

(exercise)

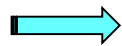
$(Y_1)_{tc} \neq 0, (Y_2)_{tc} \neq 0, (Y_1)_{tt} = (Y_2)_{tt} \neq 0$ (3 params.)
all else vanish, and take $t_\beta = 1$ for convenience

then

$$\sqrt{2} Y^{\text{SM}} = Y_1 + Y_2 \quad \text{diag. by just } V_R^u$$

but

$$-Y_1 + Y_2 \quad \text{not diag.}$$



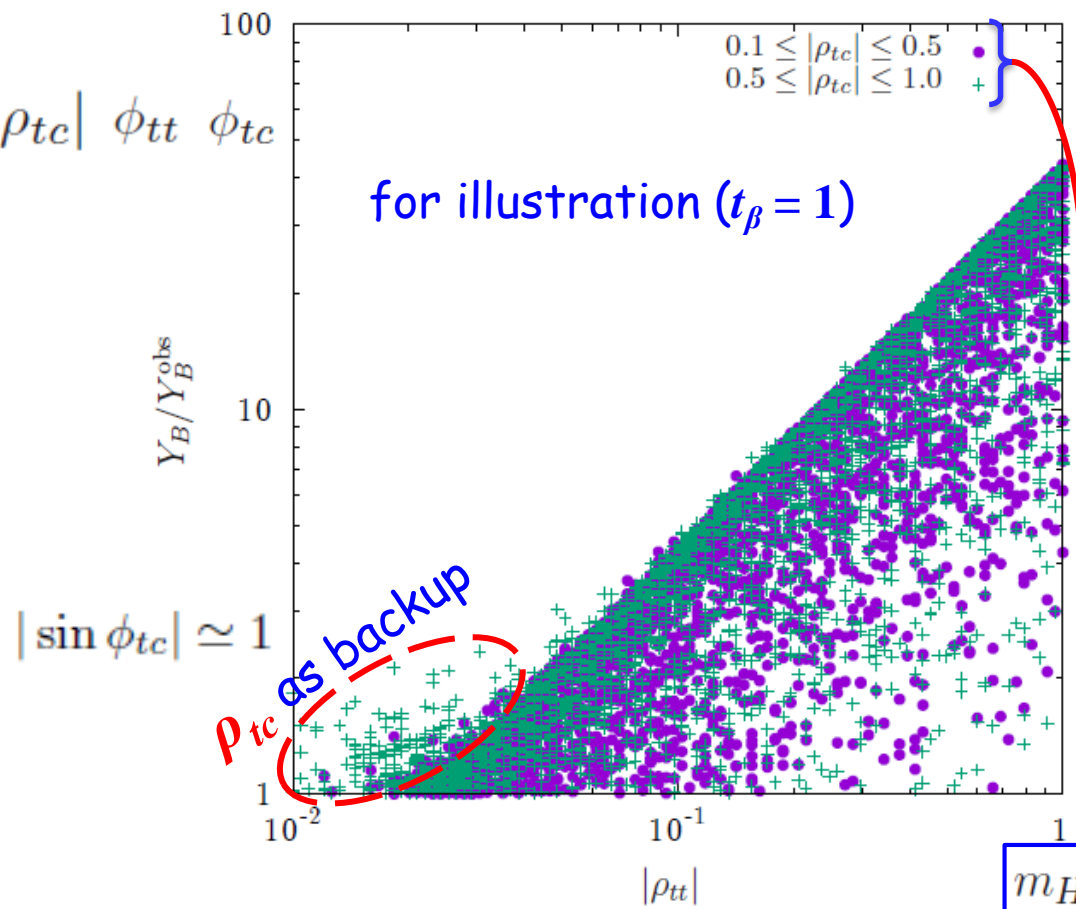
$$\text{Im} [(Y_1)_{tc} (Y_2)_{tc}^*] = -y_t \text{Im}(\rho_{tt}), \quad \rho_{ct} = 0$$

ρ_{tc} still basically free param.



Robust: Large Parameter Space for EWBG

scan over $|\rho_{tc}|$ ϕ_{tt} ϕ_{tc}



ρ_{tc}, ρ_{tt} satisfy $B_{d,s}$ mixing, $b \rightarrow sy$

no obvious diff.
 $\Rightarrow \rho_{tt}$ driven!

the charm of EWBG

$$m_H = m_A = m_{H^\pm} = 500 \text{ GeV}$$



small ρ_{tt}

$$v_C/T_C > \mathcal{O}(1)$$

$T_C = 119.2 \text{ GeV}$	$v_C = 176.7 \text{ GeV}$	$v_w = 0.4$	$\Delta\beta = 0.015$	$D_q = 8.9/T$	$D_H = 101.9/T$
$m_{t_L} = 0.59T$	$m_{t_R} = 0.62T$	$m_{c_R} = 0.50T$	$\Gamma_{qL,R} = 0.22T$	$\Gamma_B^{(s)} = 120\alpha_W^5 T$	$\Gamma_{ss} = 16\alpha_s^4 T$



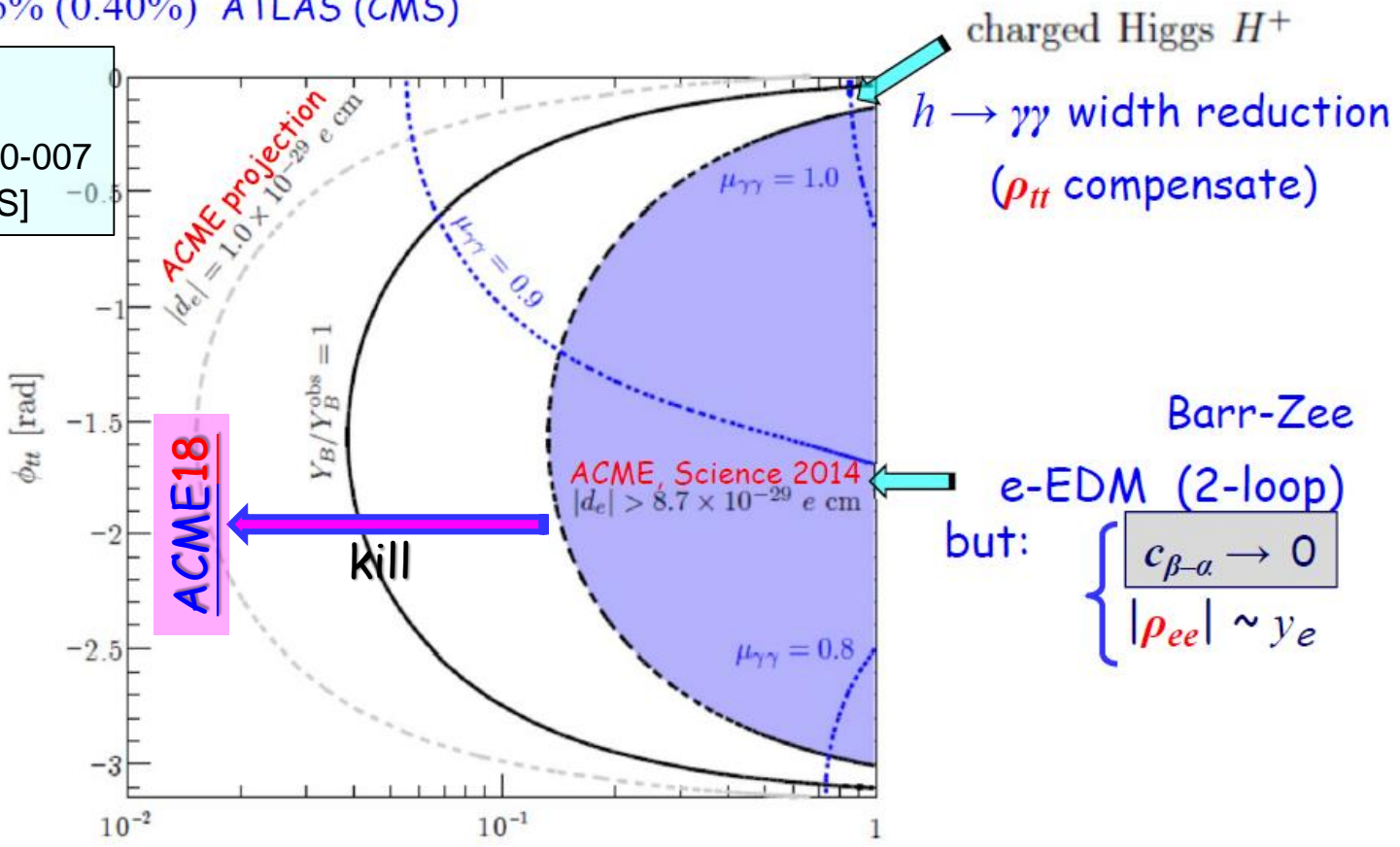
Simplified flop: $eEDM$ "Prediction"

for illustration: $t_\beta = 1$, $c_{\beta-a} = 0.1$

$\mathcal{B}(t \rightarrow ch) \simeq 0.15\%$ for $|\rho_{tc}| = 1$

vs $< 0.46\%$ (0.40%) ATLAS (CMS)

$< 0.073\%$
CMS TOP-20-007
[NTUCMS]



FHS'18 Fuyuto, WSH, Senaha, Phys. Lett. B 776 (2018) 402

the Advanced Cold Molecule Electron EDM Experiment



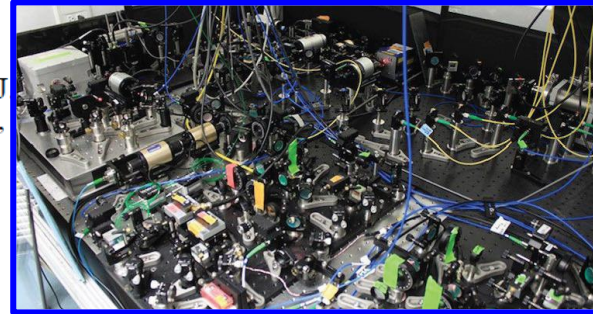
Science

AAAS

Order of Magnitude Smaller Limit on the Electric Dipole Moment of the Electron

The ACME Collaboration, J. Baron, W. C. Campbell, D. DeMille, J. M. Doyle, G. Gabrielse, Y. V. Gurevich, P. W. Hess, N. R. Hutzler, E. Kirilov, I. Kozyryev, B. R. O'Leary, C. D. Panda, M. F. Parsons, E. S. Petrik, B. Spaun, A. C. Vutha and A. D. West (December 19, 2013)

Science **343** (6168), 269-272. [doi: 10.1126/science.1248213] originally published online December 19, 2013



Editor's Summary

polar molecule thorium monoxide, we measured $d_e = (-2.1 \pm 3.7_{\text{stat}} \pm 2.5_{\text{sys}}) \times 10^{-29} e \cdot \text{cm}$. This corresponds to an **upper limit of $|d_e| < 8.7 \times 10^{-29} e \cdot \text{cm}$ with 90% confidence**, an order of magnitude

Stubbornly Spherical

JILA'17 (Cornell): $< 13 \times 10^{-29} e \text{ cm}$

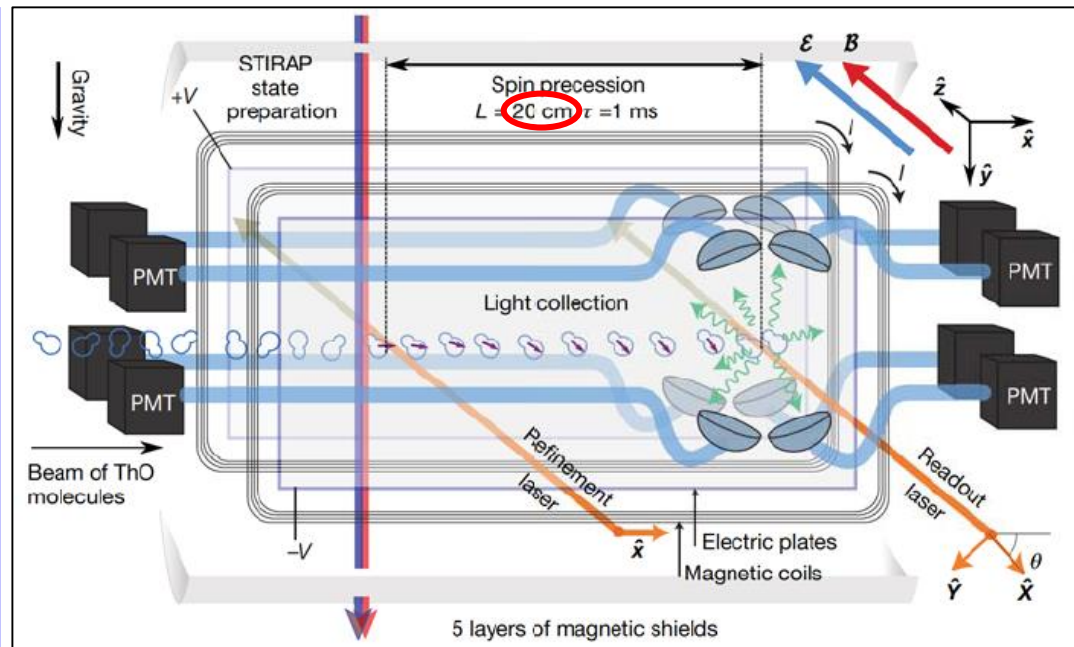


$$|d_e| < 1.1 \times 10^{-29} e \text{ cm} \quad (5)$$

at 90% confidence level. This is **8.6 times smaller** than the best previous limit, from ACME I^{1,9}. Because paramagnetic molecules are sensitive to multiple time-reversal-symmetry-violating effects³⁴, our measurement can be more generally interpreted as $\hbar\omega^{N\bar{c}} = -d_e \mathcal{E}_{\text{eff}} + W_S C_S$, where C_S is a dimensionless time-reversal-symmetry-violating **electron-nucleon coupling parameter** and $W_S = -2\pi\hbar \times 282 \text{ kHz}$ is a molecule-specific constant^{16,17,35}. For the d_e limit given above, we assume $C_S = 0$. Assuming $d_e = 0$ instead gives $|C_S| < 7.3 \times 10^{-10}$ (90% confidence level).

Because the values of d_e and C_S predicted by the standard model are many orders of magnitude below our sensitivity^{2,3}, this measurement is a background-free probe for new physics beyond the standard model. Nearly every extension of the standard model⁴⁻⁶ introduces the possibility for new particles and new time-reversal-symmetry-violating phases, ϕ_T , that can lead to measurable EDMs. Within typical extensions of the standard model, an EDM arising from new particles

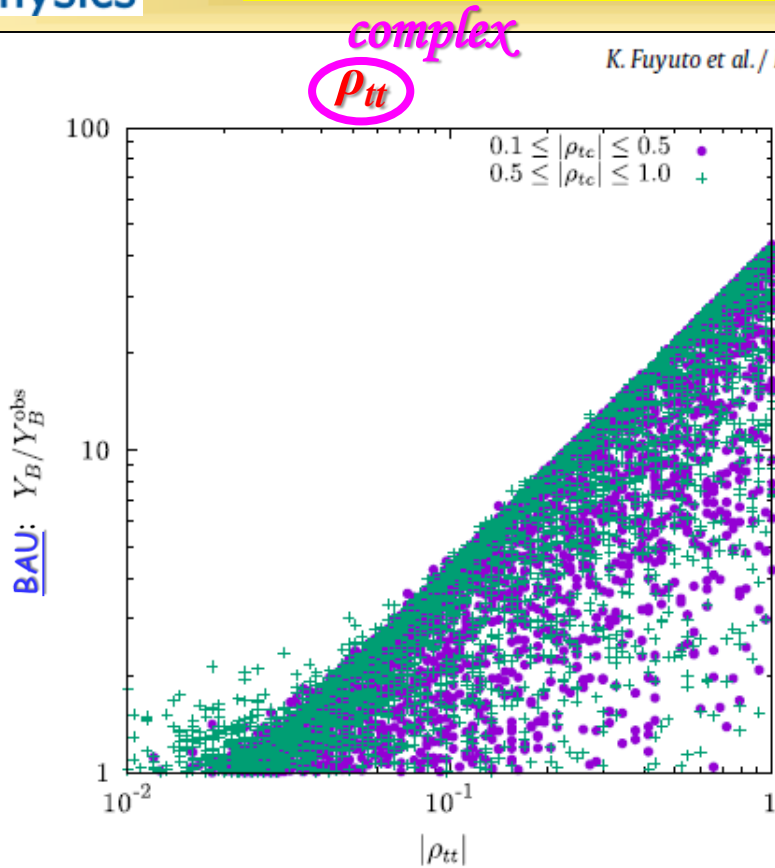
18 OCTOBER 2018 | VOL 562 | NATURE | 359



III. Under the Heavens on Earth: $eEDM$



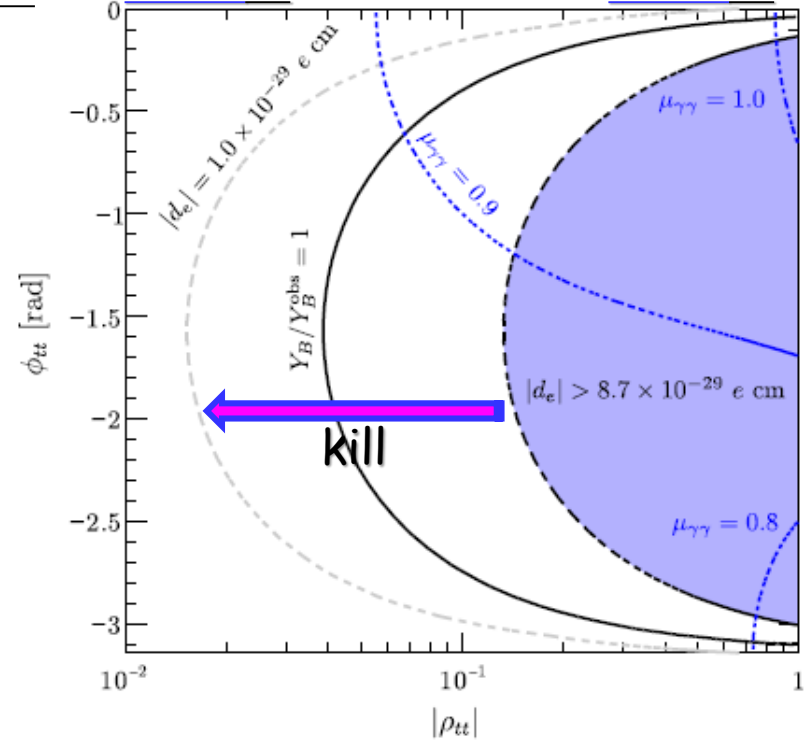
K. Fuyuto et al. / Physics Letters B 776 (2018) 402-406



FHS'18

ACME18

ACME14



EWBG



$\lambda_t \text{Im} \rho_{tt}$ robust driver

$\mathcal{O}(\lambda_t) \approx 1$

$[\rho_{tc}]$ as backup

\oplus

$eEDM: \lambda_e \text{Im} \rho_{tt}$

Ruled Out by ACME18!

Mech. to render small? Yes!



Cancellation Mechanism for d_{ThO}

to survive

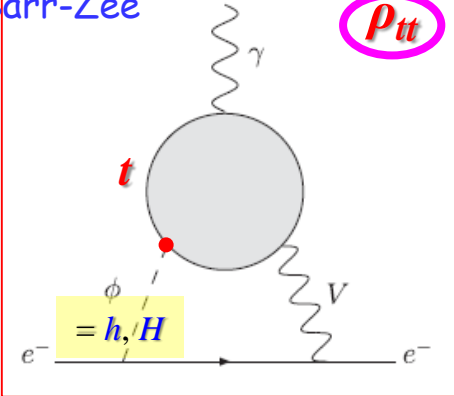
complex

dom.

$$d_e = d_e^{\phi\gamma} + d_e^{\phi Z} + d_e^{\phi W}$$

need cancellation

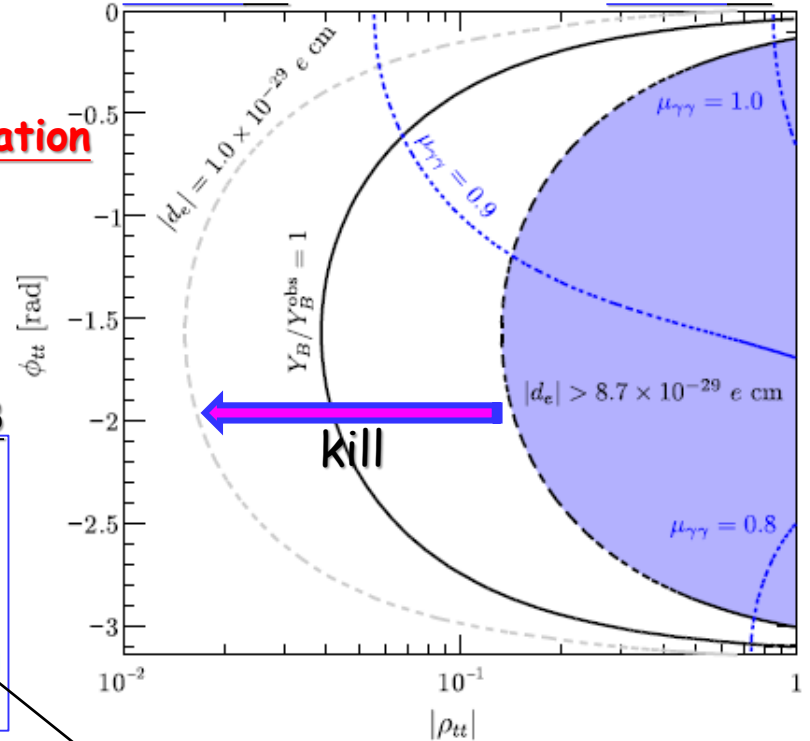
Barr-Zee



ρ_{tt}

ACME18

ACME14



FHS'18

$$\frac{(d_e^{\phi\gamma})_t}{e} = \frac{\alpha_{em} s_{2\gamma}}{12\sqrt{2}\pi^3 v} \frac{m_e}{m_t} \text{Im} \rho_{tt} \Delta g,$$

$$= -6.6 \times 10^{-29} \left(\frac{s_{2\gamma}}{0.2} \right) \left(\frac{\text{Im} \rho_{tt}}{-0.1} \right) \left(\frac{\Delta g}{0.94} \right)$$

Ruled Out

EWBG

$\lambda_t \text{Im} \rho_{tt}$ robust driver

$$\mathcal{O}(\lambda_t) \approx 1$$

$$g(m_t^2/m_h^2) - g(m_t^2/m_H^2)$$

$$g(z) \equiv \frac{1}{2} z \int_0^1 dx \frac{1}{x(1-x)-z} \ln \frac{x(1-x)}{z}$$



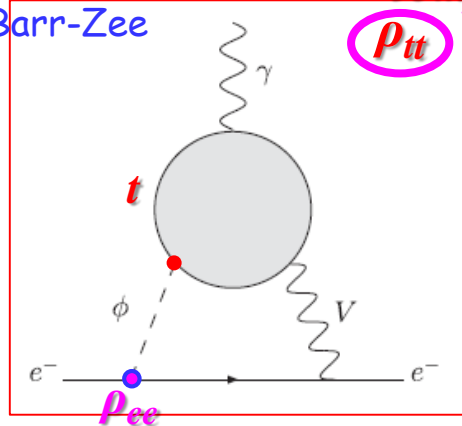
Cancellation Mechanism for d_{Th0}

to survive ACME18:

turn on ρ_{ee}

FHS'20

Barr-Zee



complex

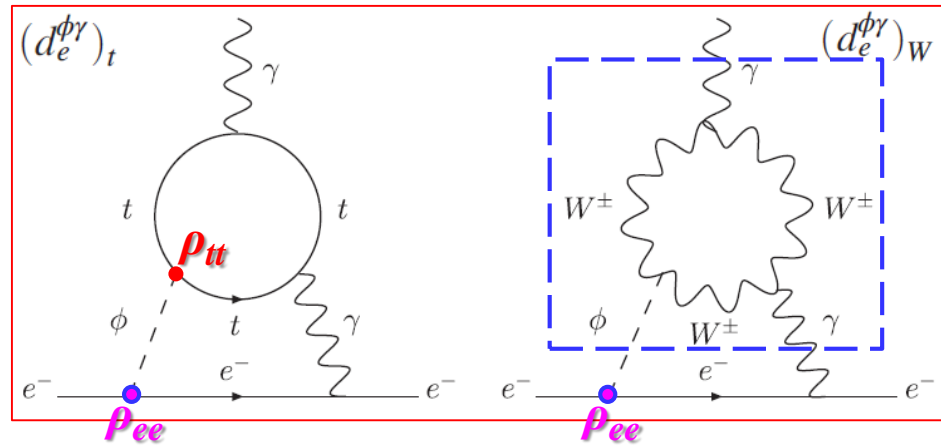
dom.

$$d_e = d_e^{\phi\gamma} + d_e^{\phi Z} + d_e^{\phi W}$$

need cancellation

N.B.

$h \rightarrow \gamma\gamma$: W -loop $>$ t -loop



h - H mixing

$$\frac{(d_e^{\phi\gamma})_{t}^{\text{mix}}}{e} = \frac{\alpha_{em} S_{2\gamma}}{12\sqrt{2}\pi^3 v} \left[\text{Im}\rho_{ee} \Delta f + \frac{m_e}{m_t} \text{Im}\rho_{tt} \Delta g \right]$$

$$\frac{(d_e^{\phi\gamma})_{W}^{\text{mix}}}{e} = -\frac{\alpha_{em} S_{2\gamma}}{64\sqrt{2}\pi^3 v} \text{Im}\rho_{ee} \Delta \mathcal{J}_W^\gamma$$

Cancel

$$\frac{\text{Im}\rho_{ee}}{\text{Im}\rho_{tt}} = c \times \frac{\lambda_e}{\lambda_t} \quad c = (16/3)\Delta g / (\Delta \mathcal{J}_W^\gamma - (16/3)\Delta f)$$

purely extr. Yuk.

$$\frac{(d_e^{\phi\gamma})_{t}^{\text{extr}}}{e} \simeq \frac{\alpha_{em}}{12\pi^3 m_t} \text{Im}(\rho_{ee}\rho_{tt}) [f(\tau_{tA}) + g(\tau_{tA})]$$

0

$[m_H \rightarrow m_A]$

$$\frac{\text{Re}\rho_{ee}}{\text{Re}\rho_{tt}} = -\frac{\text{Im}\rho_{ee}}{\text{Im}\rho_{tt}}$$

$$\left| \frac{\rho_{ee}}{\rho_{tt}} \right| = c \frac{\lambda_e}{\lambda_t}$$

w/ correlated phase



**cancel.
mech.**

$$d_{\text{ThO}} = d_e + \alpha_{\text{ThO}} C_S$$

ACME18

$$(4.3 \pm 4.0) \times 10^{-30} \text{ e cm}$$

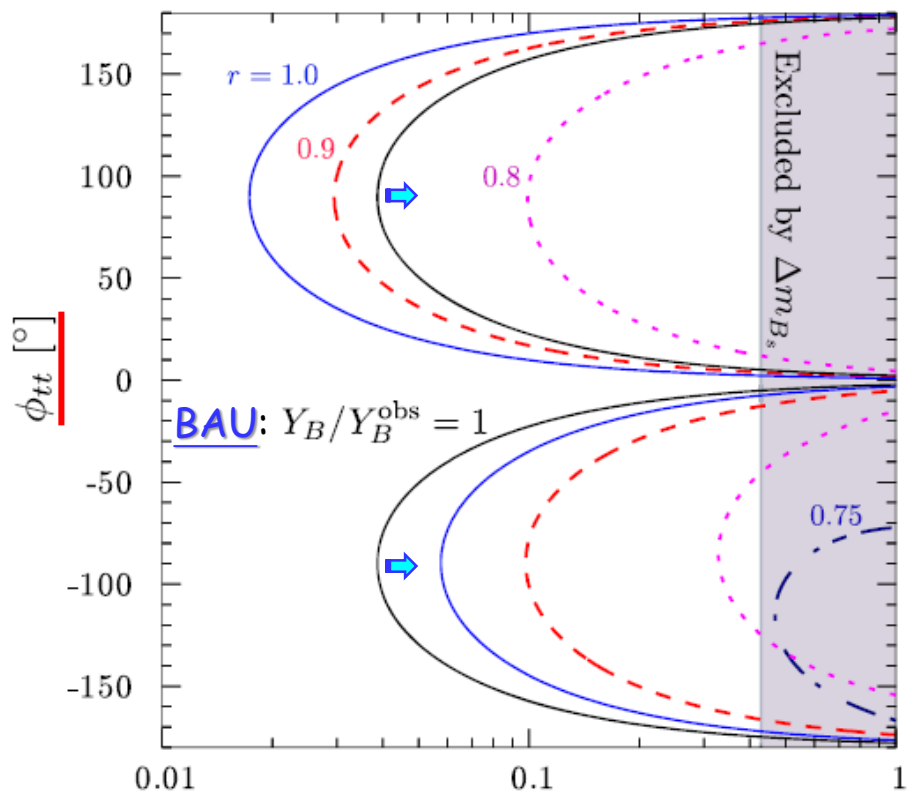
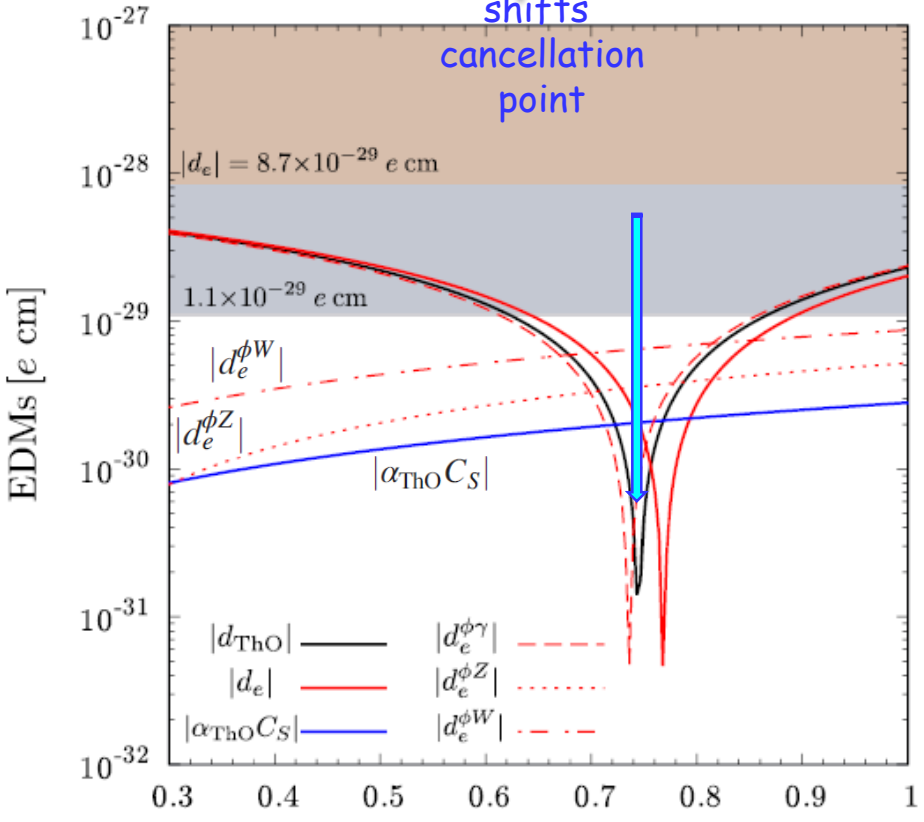
$$\alpha_{\text{ThO}} = 1.5 \times 10^{-20}$$

FHS'20

$$d_e = d_e^{\phi\gamma} + \underbrace{d_e^{\phi Z} + d_e^{\phi W}}_{\text{shifts}}$$

$$-\frac{i}{2} d_e (\bar{e} \sigma^{\mu\nu} \gamma_5 e) F_{\mu\nu}$$

$$-\frac{G_F}{\sqrt{2}} C_S (\bar{N} N) (\bar{e} i \gamma_5 e)$$



$$C_S = -2v^2 \left[6.3(C_{ue} + C_{de}) + C_{se} \frac{41 \text{ MeV}}{m_s} + C_{ce} \frac{79 \text{ MeV}}{m_c} + 0.062 \left(\frac{C_{be}}{m_b} + \frac{C_{te}}{m_t} \right) \right]$$

consistent w/ Cesarotti, Lu, Nakai, Parikh, Reece, JHEP'18

simplified
"Ansatz"

$$\frac{\text{Im } \rho_{ff}}{\text{Im } \rho_{tt}} = r \frac{\lambda_f}{\lambda_t}$$

$$\frac{\text{Re } \rho_{ff}}{\text{Re } \rho_{tt}} = -r \frac{\lambda_f}{\lambda_t}$$

Follow SM Hierarchy

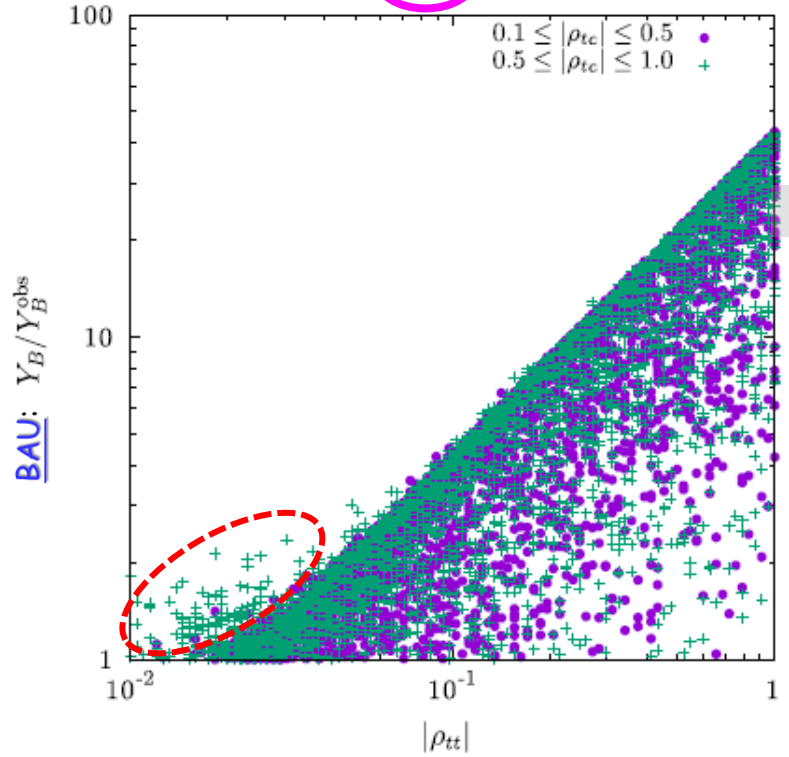
N.B. r can be f -dep.



Baryogenesis & electron EDM

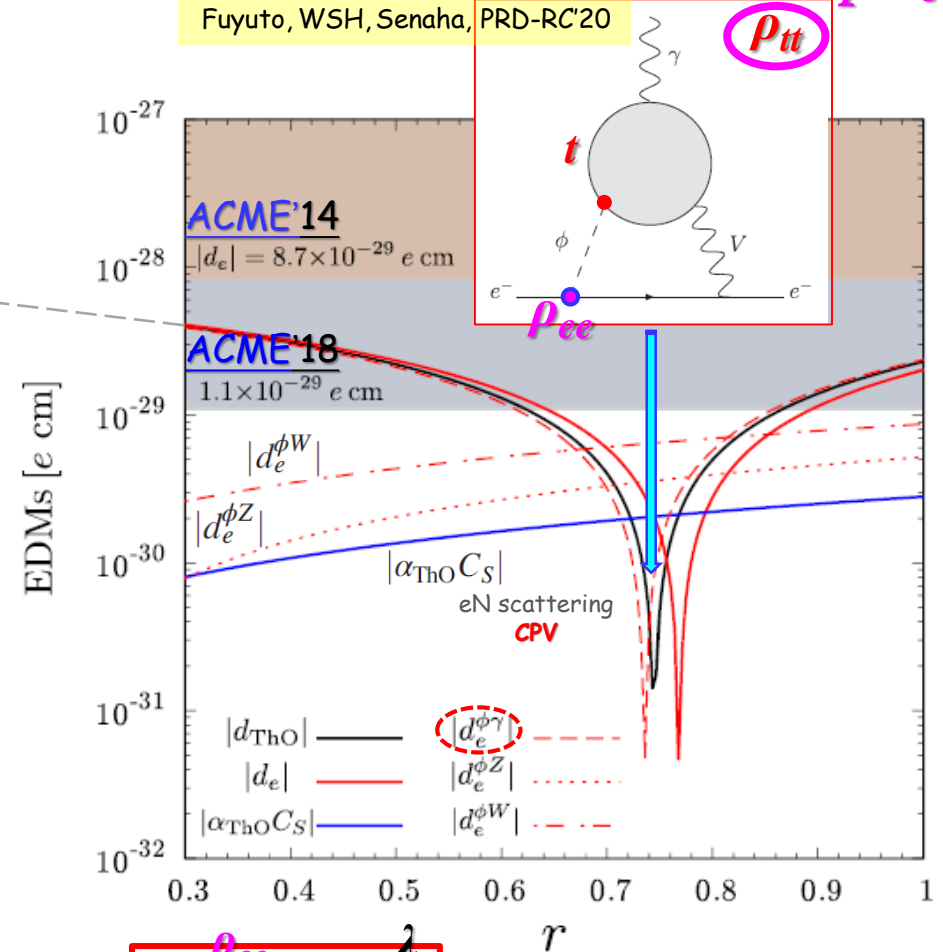
Fuyuto, WSH, Senaha, PLB'18

complex ρ_{tt}



Fuyuto, WSH, Senaha, PRD-RC'20

complex ρ_{tt}



EWBG



$\lambda_t \text{Im} \rho_{tt}$ robust driver

$$\mathcal{O}(\lambda_t) \approx 1$$

$[\rho_{tc}$ as backup

simplified "Ansatz"

$$\frac{\text{Im} \rho_{ff}}{\text{Im} \rho_{tt}} = r \frac{\lambda_f}{\lambda_t}$$

$$\frac{\text{Re} \rho_{ff}}{\text{Re} \rho_{tt}} = -r \frac{\lambda_f}{\lambda_t}$$

Follow SM Hierarchy!

N.B. r depend on loop functions



complex

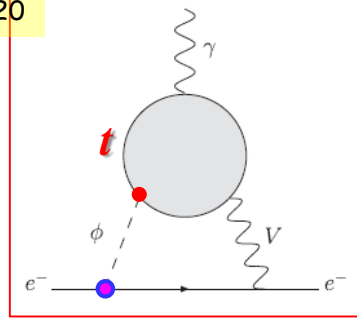
complex



P_{tt} Baryogenesis & electron EDM P_{ee}

Fuyuto, WSH, Senaha, PLB'18

Fuyuto, WSH, Senaha, PRD-RC'20



O Lord, our Lord,
How Majestic is Thy Name
in all the Earth,
Who have set Thy Splendor
above the Heavens !

Psalm 8:1 (of David)

EWBG



$\lambda_t \text{Im} \rho_{tt}$ robust driver

$\mathcal{O}(\lambda_t) \approx 1$

$[\rho_{te}]$ as backup

simplified
"Ansatz"

$$\frac{\text{Im} \rho_{ff}}{\text{Im} \rho_{tt}} = r \frac{\lambda_f}{\lambda_t}$$

$$\frac{\text{Re} \rho_{ff}}{\text{Re} \rho_{tt}} = -r \frac{\lambda_f}{\lambda_t}$$

Follow SM Hierarchy!

N.B. r depend on loop functions



complex

complex



ρ_{tt} Baryogenesis & electron EDM ρ_{ee}

Fuyuto, WSH, Senaha, PLB'18

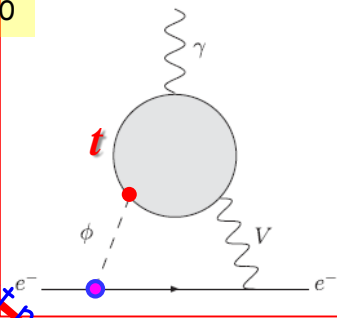
Fuyuto, WSH, Senaha, PRD-RC'20

We are Probing Extra Yukawa Couplings:

$$t \rightarrow ch; h \rightarrow \mu\tau$$

$$\rho_{tc}; \rho_{\mu\tau}$$

$$*c_\gamma$$



Alignment protection:
 c_γ small (Emergent)

Glashow-Weinberg'77: Absence of 2nd "Yukawa"!

G-W Knew:

$$\begin{aligned} m_e &\ll m_\mu \ll m_\tau, \\ m_d &\ll m_s \ll m_b, \\ m_u &\ll m_c \ll m_t, \end{aligned}$$

Did not expect:

$$m_t/m_b \gg 1,$$

1986+

Totally Out of Whack:

ca. 1983

$$|V_{ub}|^2 \ll |V_{cb}|^2 \ll |V_{us}|^2 \ll |V_{tb}|^2 \approx 1,$$

Mass-Mixing Hierarchy

→ Nature's Design

EWBG

$$\lambda_t \text{Im} \rho_{tt} \text{ robust driver}$$

$$\mathcal{O}(\lambda_t) \approx 1$$

$[\rho_{tc}]$ as backup

simplified "Ansatz"

$$\begin{aligned} \frac{\text{Im} \rho_{ff}}{\text{Im} \rho_{tt}} &= r \frac{\lambda_f}{\lambda_t} \\ \frac{\text{Re} \rho_{ff}}{\text{Re} \rho_{tt}} &= -r \frac{\lambda_f}{\lambda_t} \end{aligned}$$

Follow SM Hierarchy!

N.B. r depend on loop functions



On “the Heavens and the Earth”

P_{ee}

In context of EWBG driven by an extra top Yukawa coupling, the impressive ACME18 bound suggests an extra electron Yukawa coupling that works in concert to give **exquisite cancellation** among dangerous diagrams. The cancellation mechanism calls for the extra Yukawas to *echo* the hierarchical pattern of SM Yukawa couplings.

P_{tt}

- ACME14 was confirmed by JILA17 using HfF⁺, i.e. different approach. It is good that “JILA is chasing Harvard/Yale(/Northwestern).” Personally, I don't think 10⁻²⁹ ecm is finished business.
- Amusing: The largest diagonal *extra* Yukawa *P_{tt}* drives B.A.U., in concert w/ smallest diagonal *extra* Yukawa *P_{ee}* to generate *eEDM*, that might be revealed soon by very-low-energy ultraprecision probes. For me, I think 10⁻²⁹ - 10⁻³⁰ ecm is just fabulous. **Godspeed success!**



extra Yukawas reflect SM Yukawa pattern

$$\rho_{ii} \lesssim \mathcal{O}(\lambda_i); \quad \rho_{1i} \lesssim \mathcal{O}(\lambda_1); \quad \rho_{3j} \lesssim \mathcal{O}(\lambda_3) \quad (j \neq 1)$$

IV. Phenomenological Consequences

H/A/H⁺ Search & Flavor Frontier



Leading Search Modes at the LHC

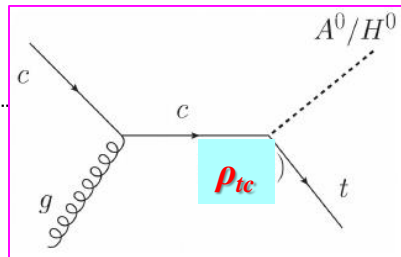


G2HDM

Sub-TeV Spectrum

GeV

Search Zone



Search at ATLAS/CMS started

$cg \rightarrow tH/A \xrightarrow{\rho_{tc}} tt(\bar{c})$

Kohda, Modak, WSH, PLB'18

Same-Sign Top

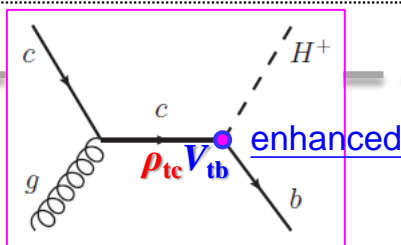
$\xrightarrow{\rho_{tt}} ttt(\bar{c})$

Triple-Top (High Lumi LHC; higher mass, more exquisite, tiny SM)

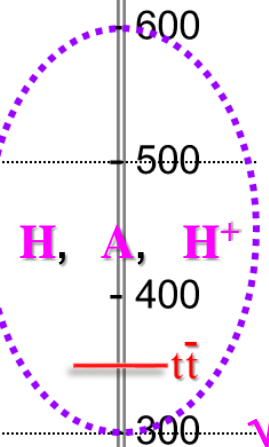
$cg \rightarrow bH^+ \rightarrow bt(\bar{b})$

Top w/ two p_T b-jets (H^+)

Ghosh, WSH, Modak, 1912.10613 (PRL'20)



v.e.v.



HKM, PLB'18

ν
neutrino

e, μ, τ
lepton

d, s, b
down-type

u, c
up-type

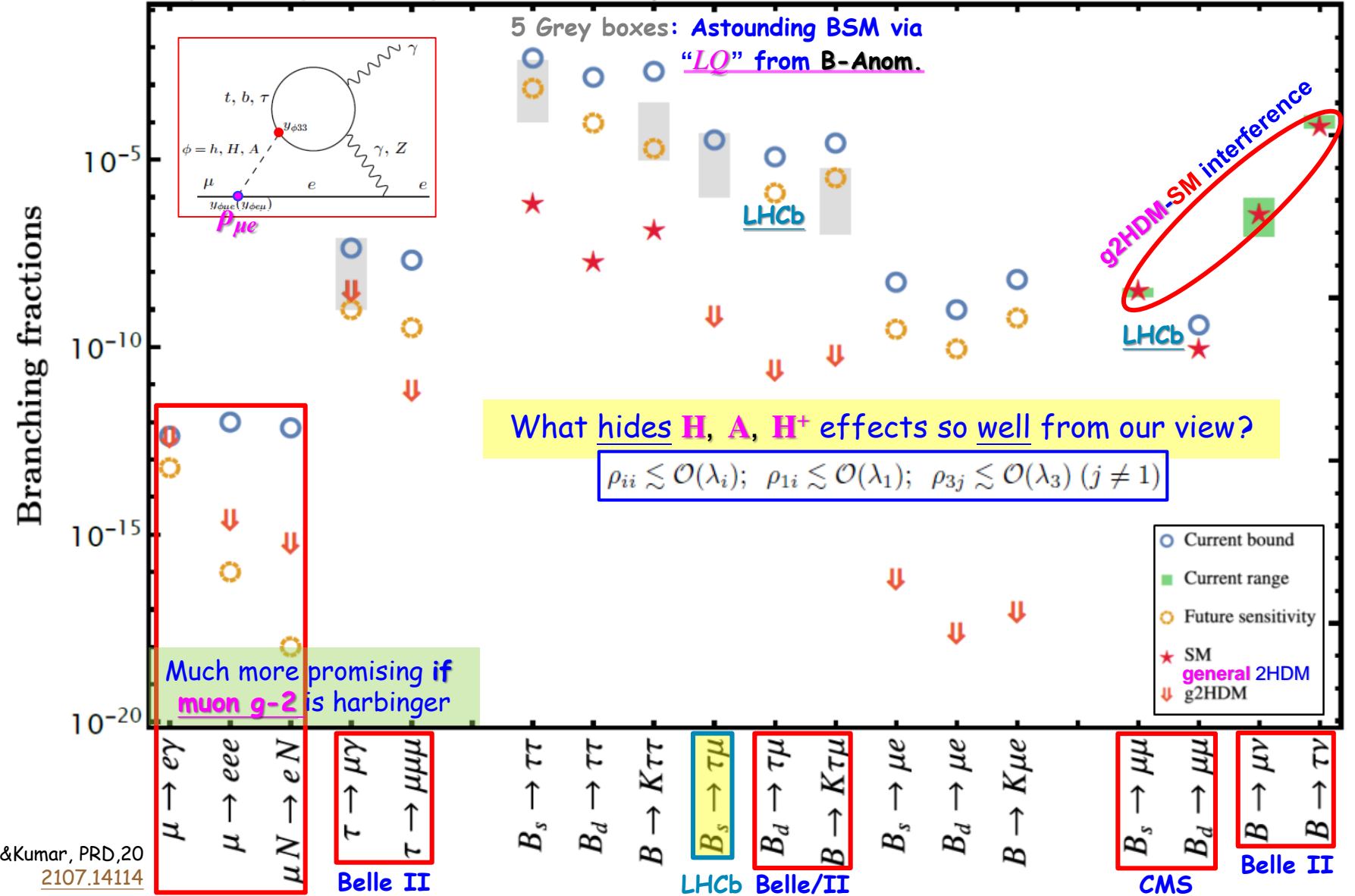
γ, g
vector

scalar



Glimpse of coming New Flavor Era

μ & τ FV (Flav. Viol.) in B decay



WSH&Kumar, PRD,20 2107.14114



Our Life & Times

$h(125)$ ✓ *New Physics* ✗

High Scale SUSY?!

Where is SUSY?

unconventional-Conventional
Road Not Taken

Extra Higgs Doublet w/
Extra Weinberg Couplings
& Extra Quartic Couplings

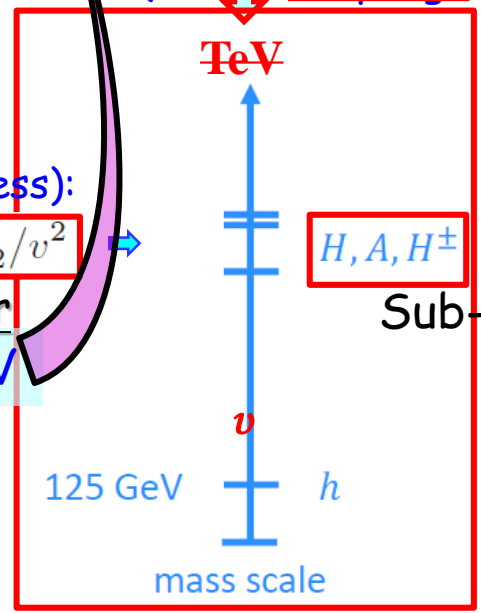
1stEWPT

Dim'less Quartics $\mathcal{O}(1)$ (Naturalness):

η_i with $i = 1-7; \mu_{22}^2/v^2$

100 TeV pp collider

→ Landau Pole ~ 10-20 TeV





IV'. Pheno Consequences: one-loop muon g-2

• $gg \xrightarrow{\rho_{tt}} H, A \xrightarrow{\rho_{\mu\tau}} \mu\tau$: stringent bound on $\rho_{tt} \rho_{\mu\tau}$, so could appear soon!
 $\rho_{tt} > 0.1$ can still drive EWBG.

• $gg \xrightarrow{\rho_{tt}} H, A \xrightarrow{\rho_{tc}} tc$: ρ_{tc} can dilute the above
 \rightarrow $cg \rightarrow bH^+ \rightarrow \mu\tau bW^+, tcbW^+$ fancy signatures.

[WSH, Jain, Kao, Kumar, Modak, 2105.11315](#)

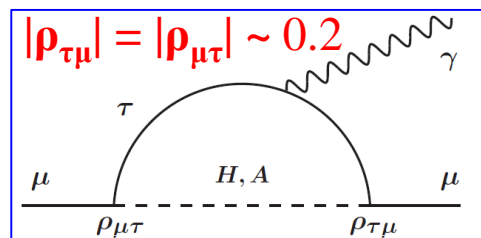
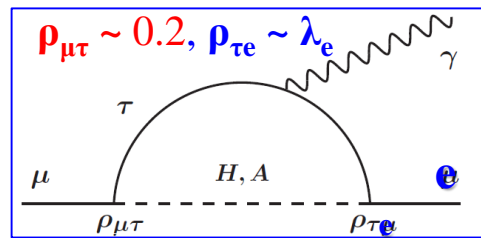
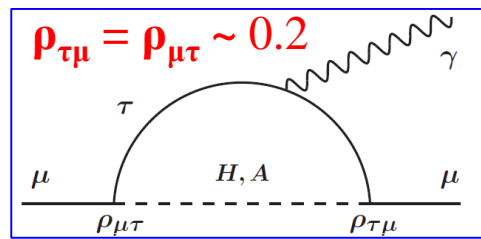
• Revival of muon physics:
 - MEG II discovery plausible (with $\rho_{\tau e} \sim \lambda_e$)
 - follow-up by $\mu N \rightarrow eN$, can even probe ρ_{qq} !
 - $\tau \rightarrow \mu\gamma$: probe $\rho_{\tau\tau} \sim \lambda_\tau$! / $\tau \rightarrow 3\mu$: probe $\rho_{\mu\mu} \sim \lambda_\mu$!

[WSH, Kumar, 2107.14114](#)

• **μ EDM**: Same one-loop diagram, complex $\rho_{\tau\mu} \rho_{\mu\tau}$
CPV \rightarrow Possibly discoverable at PSI with planned sensitivity!

[WSH, Kumar, 2109.08936](#)

$6 \times 10^{-23} \text{ e cm}$



N.B. This one-loop muon g-2 would make Nature appear “whimsical”.



V. Summary

the A and the Ω

I have told you up front:

$H^0, A^0, H^\pm \sim 500 \text{ GeV}$
can generate **B.A.U.**
accommodate $e\text{EDM}$

CAN
Verify at LHC.
and FPCP Probes !

← Fantastic!!

Decadal Mission:

Find the extra H, A, H^\pm bosons and crack the *Flavor* code!

Go CMS & Belle II (and others) !

& Lattice



Thank you!





Soaring to the Starry Heavens



Enough **CPV**
for **BAU**?



天

CPV4**BAU**

地



Belle experimentalist: KM got Nobel, but the KM **CPV phase** is short of **Baryogenesis** by 10^{-10} !

