

Keiko Nagao
(Okayama Univ. of Sci., Japan)

長尾 桂子
(岡山理科大学)



Discriminating WIMP Mass and Anisotropy with Directional Detector

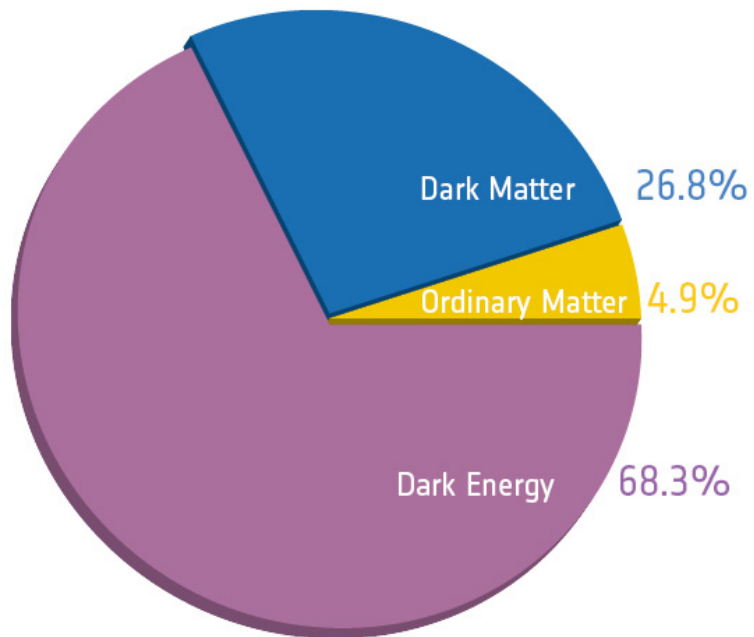
based on arXiv:1707.05523 (accdepted to Physics of the Dark Universe);
KN, R. Yakabe (Kobe Univ.), T. Naka (Nagoya Univ., Toho Univ.), K. Miuchi (Kobe Univ.)
work in progress

KN, T. Ikeda (Kobe Univ.), R. Yakabe, T. Naka, K. Miuchi

Outline

1. Introduction
2. Anisotropy of velocity distribution
3. Numerical Results
4. Summary

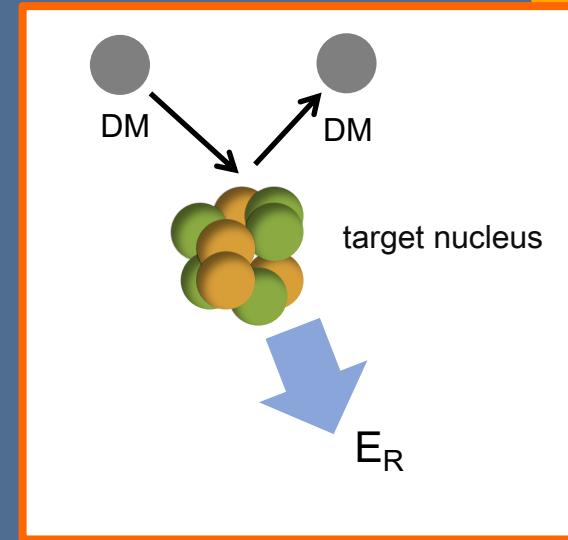
Dark Matter



- Weakly Interacting Massive Particles (WIMPs)
- Axions
- Primordial black holes
- Modified Gravity
-

Direct Detection

- Detect recoil energy of DM-target scattering



Underground facilities (a partial list)

It has been proven that underground facilities are very important for varieties of science!
For scientific reasons, It would be very nice if there is (at least) one in the Southern hemisphere...



Constraint from Direct Detection

$$R \propto N_T N_\chi f(\vec{v}) \langle v \rangle \sigma$$

$$\Rightarrow \frac{dR}{dE_R} = \frac{N_T \rho_0}{m_\chi} \int^{v_{\max}} d\vec{v} f(\vec{v}) |\vec{v}| \frac{d\sigma(\vec{v})}{dE_R}$$

R Event rate

N_T # of target particles

$N_\chi = \frac{\rho_0}{m_\chi}$ # of WIMP

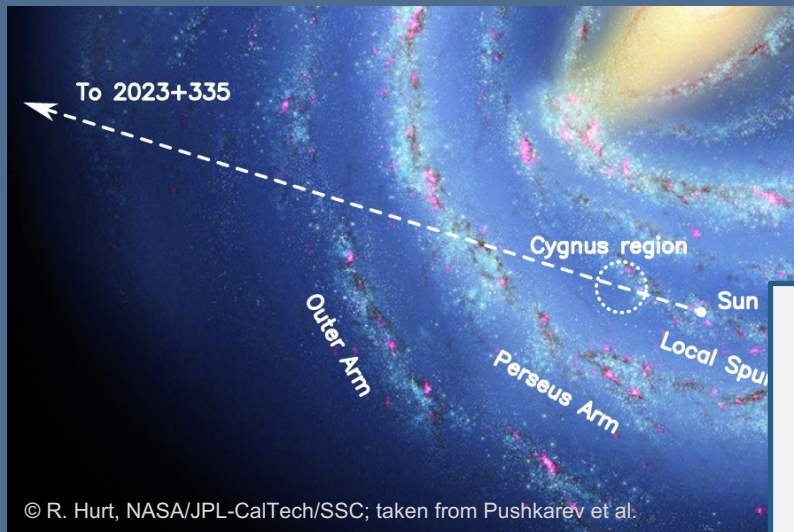
$f(\vec{v})$ Velocity distribution

$\langle v \rangle$ Averaged WIMP velocity

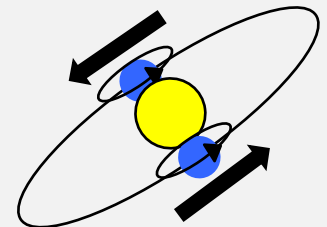
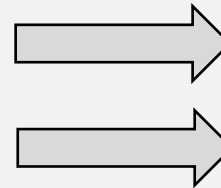
σ Cross section for
DM-nucleus scattering

Directional Detection

- detect not only the recoil energy but also direction where DM comes from.



DM wind



Directional Searches



talk by J. Mardon
in Astroparticle phys

DMTPC

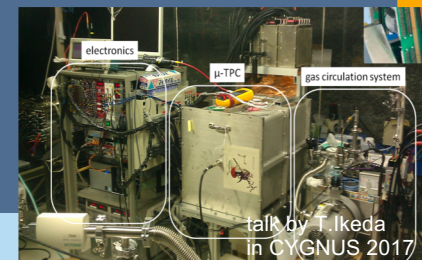
DRIFT



talk by N. Spooner
in CYGNUS 2017

MIMAC

NEWAGE



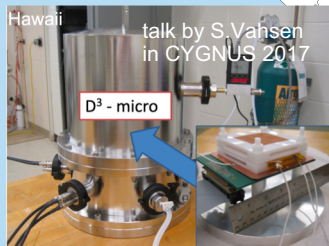
talk by T. Ikeda
in CYGNUS 2017

NEWSdm



talk by A. Um
in CYGNUS 2017

D³



Hawaii
talk by S. Vahsen
in CYGNUS 2017

D³ - micro

Gas Detector
(CF₄, CS₂, CHF₃)
SD cross section

Solid Detector
(Ag, Br, C,...)
SI cross section

(not complete list)

Outline

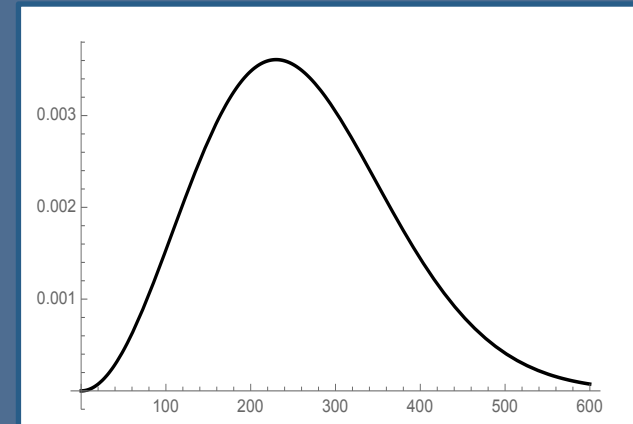
1. Introduction
2. Anisotropy of velocity distribution
3. Numerical Results
4. Summary

Standard velocity distribution

■ Maxwell-Boltzmann (MB) distribution

$$\frac{dR}{dE_R} = \frac{N_T \rho_0}{m_\chi} \int^{v_{\max}} d\vec{v} f(\vec{v}) |\vec{v}| \frac{d\sigma(\vec{v})}{dE_R}$$

$$f(v) = \frac{1}{(\pi v_0^2)^{3/2}} e^{-(v+v_E)^2/v_0^2}$$



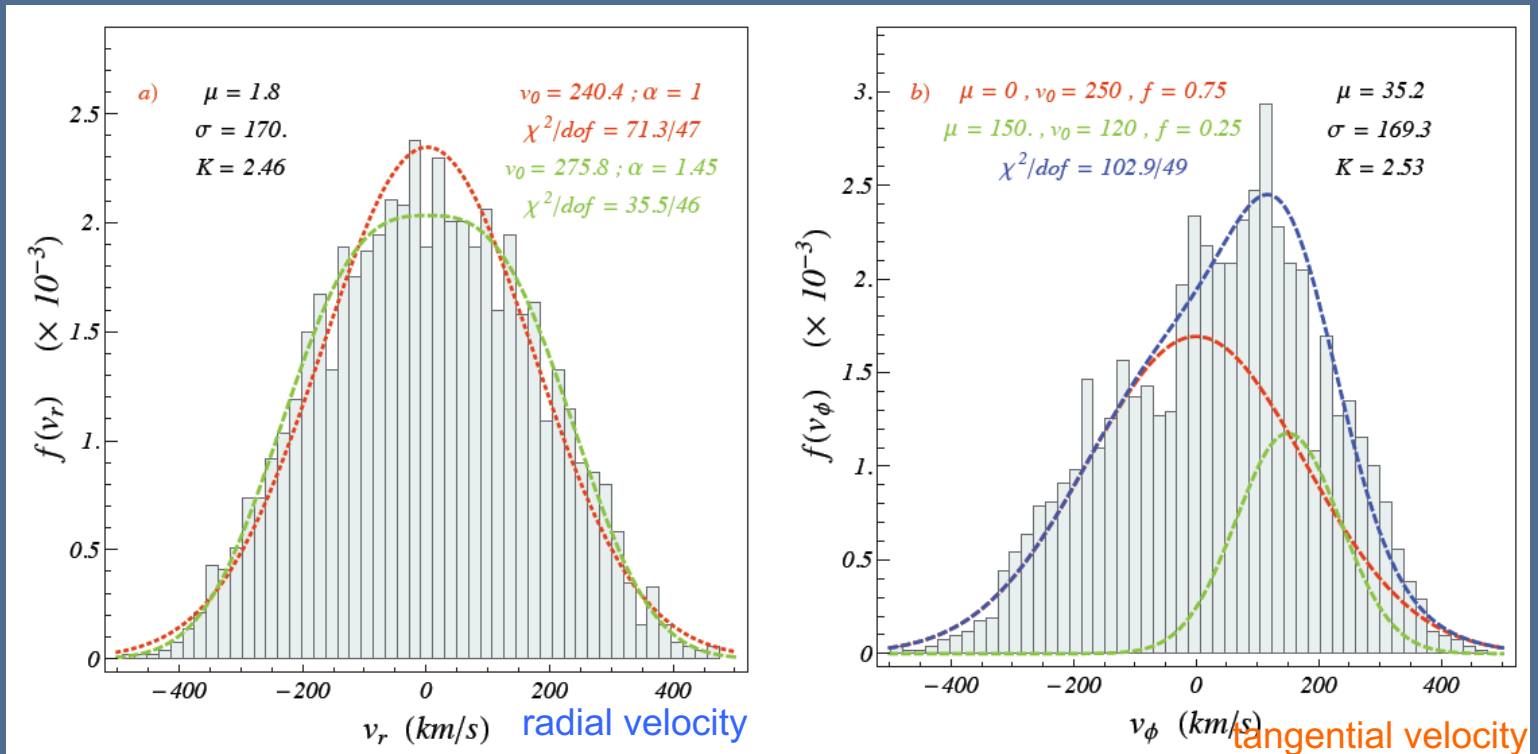
- isotropic MB distribution is commonly supposed in direct detections

Simulation including baryons and gas

the Galaxy



the Solar system

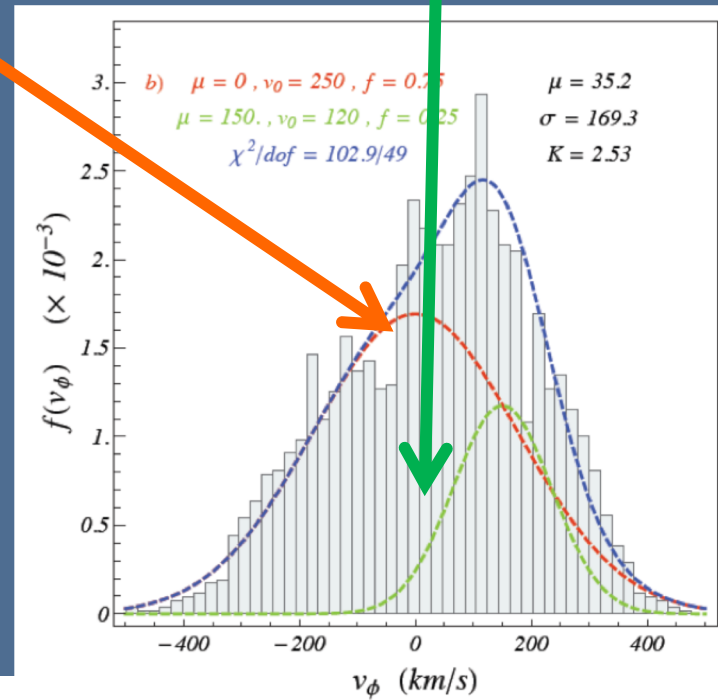


Ling, Nezri, Athanassoula & Teyssier (2009)
cf. David R. Law (2009) ...

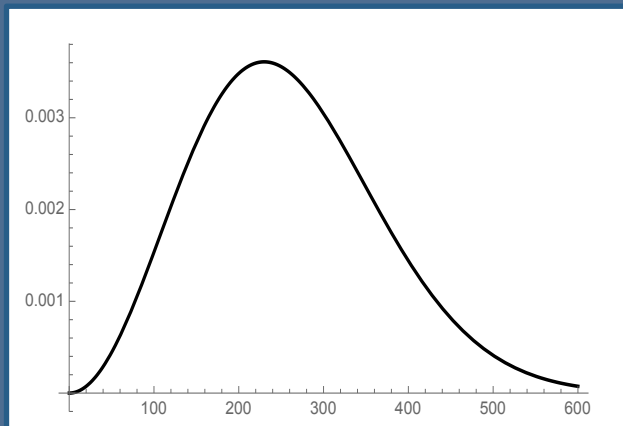
Anisotropic component

$$f(v_\phi) = \underbrace{\frac{1-r}{N(v_{0,\text{iso.}})} \exp[-v^2/v_{0,\text{iso.}}^2]}_{\text{isotropic}} + \underbrace{\frac{r}{N(v_{0,\text{ani.}})} \exp[-(v-\mu)^2/v_{0,\text{ani.}}^2]}_{\text{anisotropic}}$$

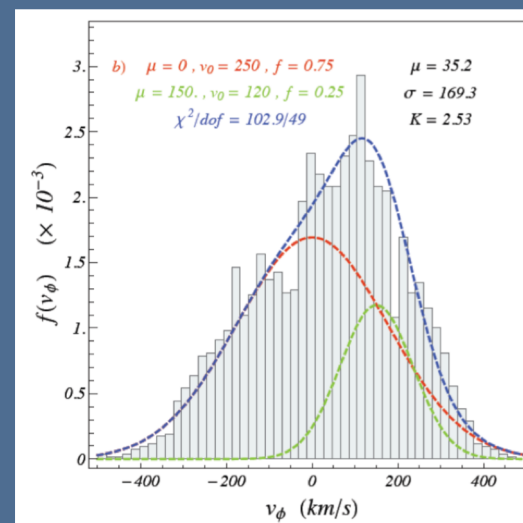
- Tangential velocity
 - anisotropy parameter r
 - $r=0.25$ is suggested by simulation



Simplified Goal



OR



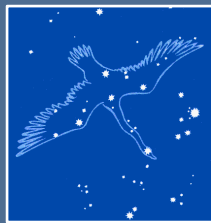
AND M_{DM}

Outline

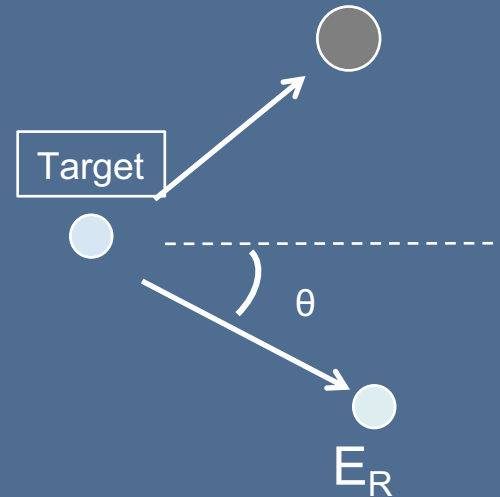
1. Introduction
2. Anisotropy of velocity distribution
3. Numerical Results
4. Summary

cf.
Ben Morgan, Anne M. Green, Neil J. C. Spooner (2004)
Ole Host, Steen H Hansen (2007)

Numerical Simulation of Scattering



DM wind $f(v)$

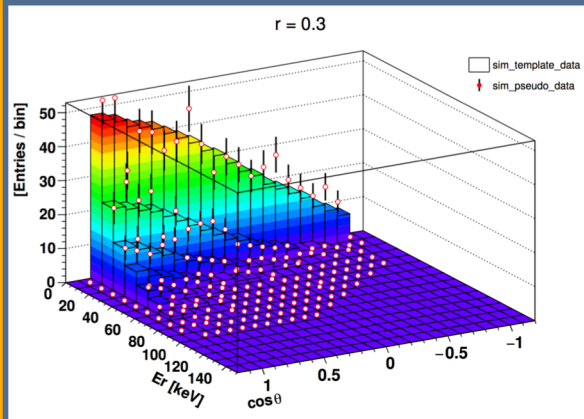


- Monte Carlo simulation of scattering **supposing** $f(v)$
 - E_R and θ are obtained
 - Elastic scattering, No BG, Perfect resolution
 - Target : F (light) / Ag (heavy)

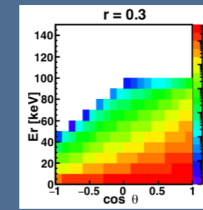
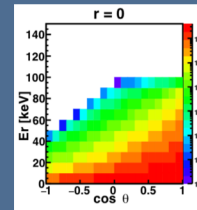
Strategy for discrimination



ideal **“template”**
Many Data
(#10⁸)



“pseudo-experimental” data
Fewer Data
(#10³-10⁴)



Which template is
more similar
to pseudo-exp?

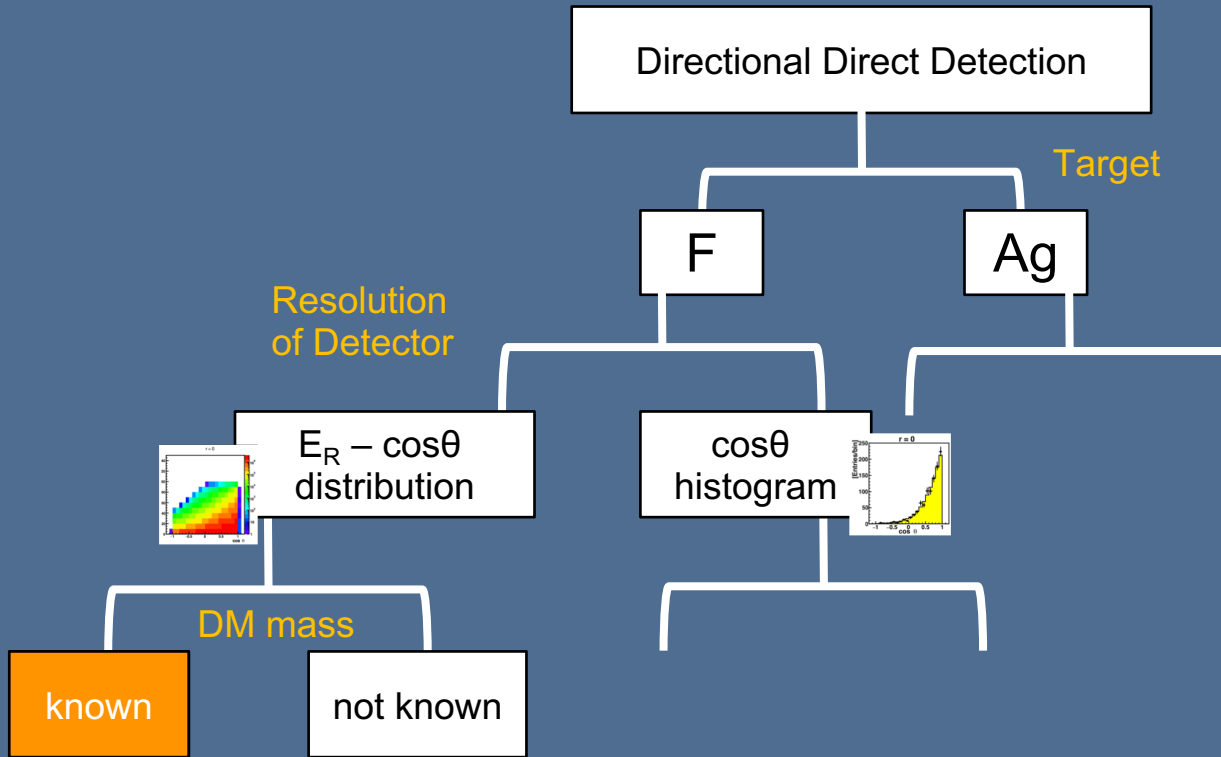
□ Likelihood estimation

$$\mathcal{L} = \prod_{\text{bins}} P(r \mid \text{pseudo, template})$$

□ χ^2 test

$$\chi^2 = \sum_{\text{bins}} \frac{(\text{pseudo} - \text{template})^2}{\text{pseudo}}$$

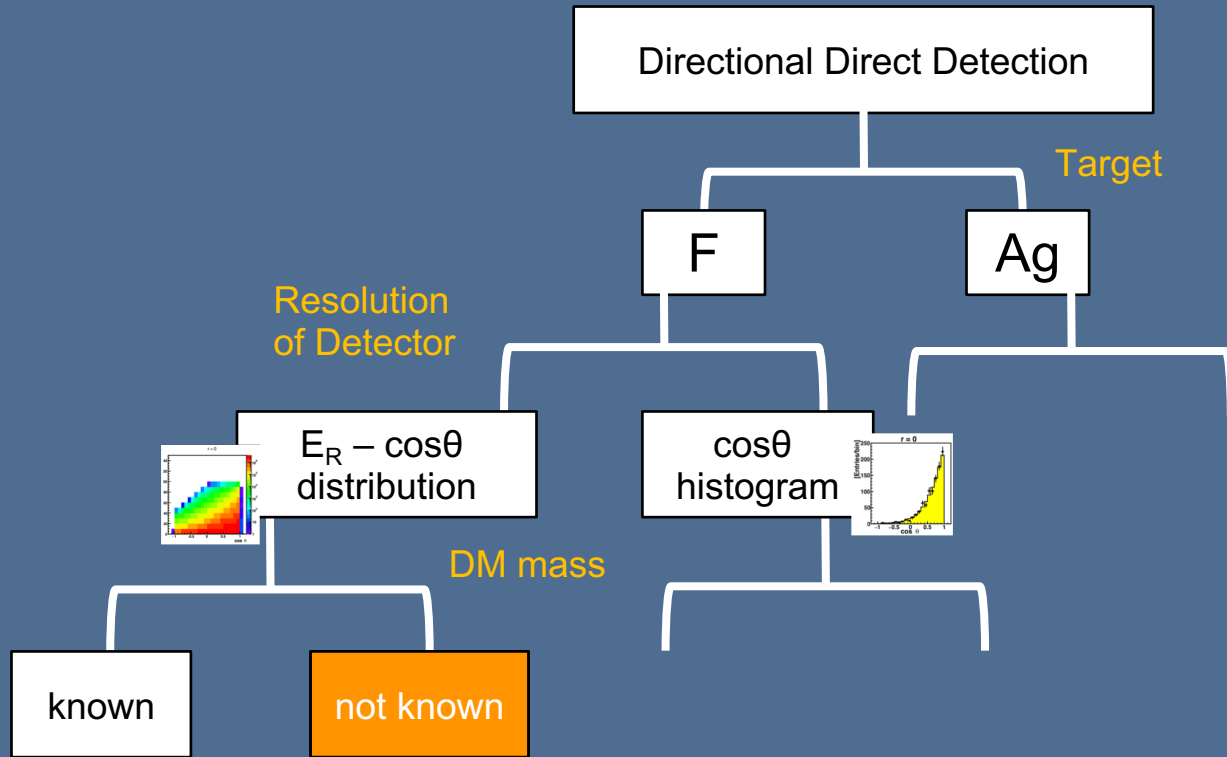
Summary of Branches



6×10^3 (ER-angle) / 5×10^3 (angle only) for F
 6×10^4 (ER-angle) / 2×10^4 (angle only) for Ag.

(arXiv:1707.05523)

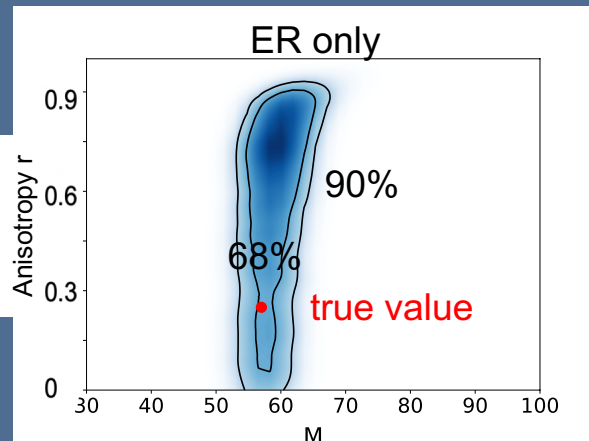
Cases



What if M_{DM} is not known?

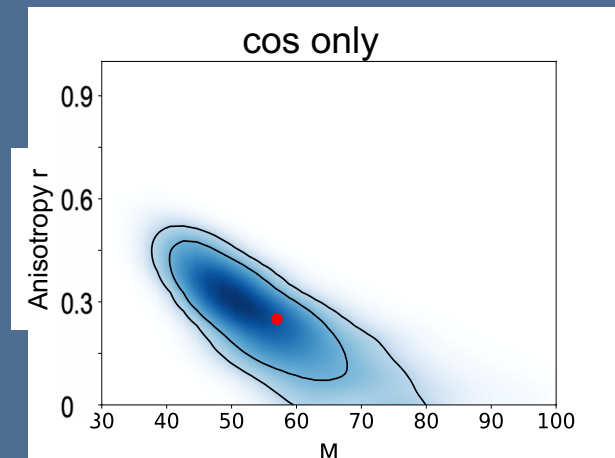
□ Likelihood method

E_{thr}=50keV (F)
 M_{dm} =60GeV
#event: 1000



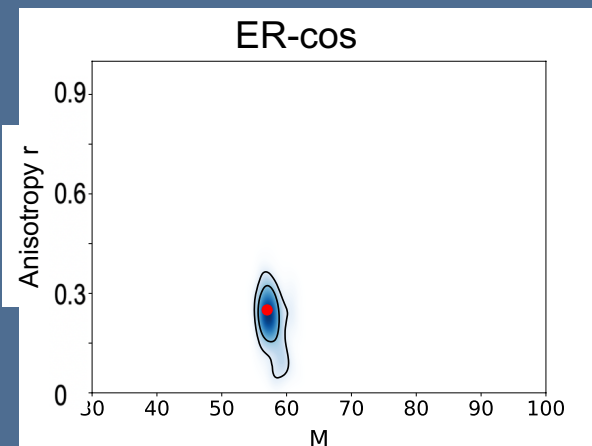
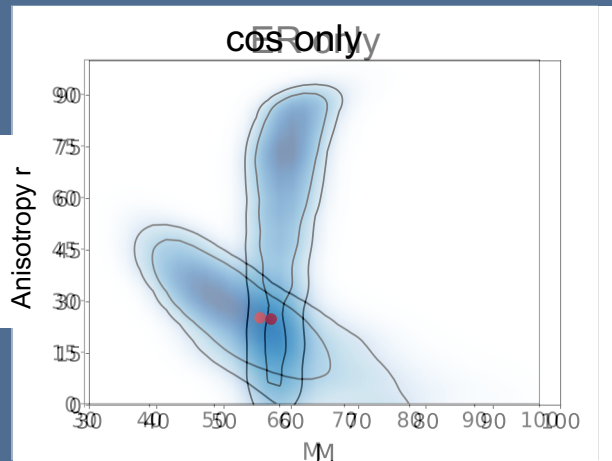
□ Anisotropy is not discriminated only by E_R .

□ Constraint for mass by directionality histogram is not so strictly.



What if M_{DM} is not known?

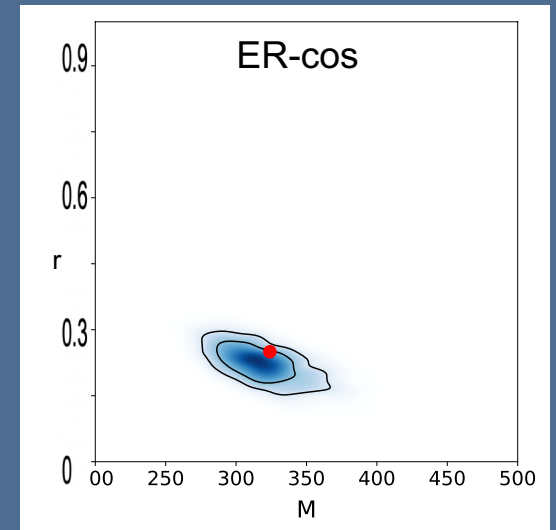
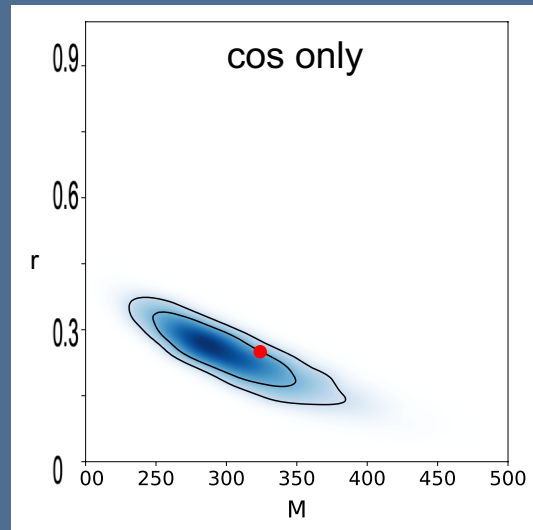
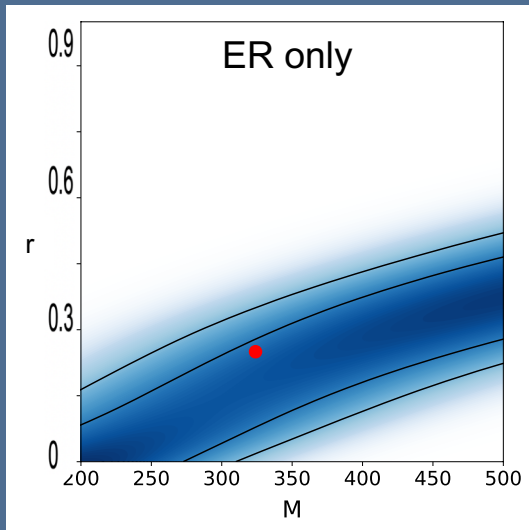
□ Likelihood method



E_{thr}=50keV (F)
M_{dm}=60GeV
#event: 1000

- We need both E_R and directional information to give constraint for both anisotropy and mass at the same time.

What if M_{DM} is not known? (Ag)



□ Likelihood method

E_{thr}=50keV (Ag)
 M_{dm} =300GeV
#event: 10000

Summary

- Possibility to figure out DM mass and anisotropy of DM distribution is discussed.
- If DM mass is known by other searches, we can discuss the anisotropy once $O(10^3-10^4)$ event is obtained in directional detection.
- Even if M_{DM} is not known, once both E_R and angular information are obtained we can give constraints for M_{DM} and distribution.

Thank you for your attention.