

Non-Standard Neutrino Interactions: From Oscillations to Colliders

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NEUTRINOS: CORNERSTONE OF THE SM

At low energies, Fermi's contact interactions:

$$-\mathcal{L}_{eff}^{cc} = \frac{G_F}{\sqrt{2}} J_W^\mu J_{W\mu}^\dagger, \quad -\mathcal{L}_{eff}^{NC} = \frac{G_F}{\sqrt{2}} J_Z^\mu J_{Z\mu}, \quad \frac{G_F}{\sqrt{2}} \simeq \frac{g^2}{8M_W^2} = \frac{1}{2\nu^2}$$

$$J_W^{\mu\dagger} = (\bar{\nu}_e \bar{\nu}_\mu \bar{\nu}_\tau) \gamma^\mu (1 - \gamma^5) V_\ell \overset{\text{PMNS}}{\begin{pmatrix} e^- \\ \mu^- \\ \tau^- \end{pmatrix}} + (\bar{u} \bar{c} \bar{t}) \gamma^\mu (1 - \gamma^5) V_q \overset{\text{CKM}}{\begin{pmatrix} d \\ s \\ b \end{pmatrix}}$$

$$J_Z^\mu = \sum_m [\bar{u}_{mL} \gamma^\mu u_{mL} - \bar{d}_{mL} \gamma^\mu d_{mL} + \bar{\nu}_{mL} \gamma^\mu \nu_{mL} - \bar{e}_{mL} \gamma^\mu e_{mL}]$$

At and above the EW scale:

$$\mathcal{L} = -\frac{g}{2\sqrt{2}} \left(J_W^\mu W_\mu^- + J_W^{\mu\dagger} W_\mu^+ \right) - \frac{g}{2 \cos \theta_W} J_Z^\mu Z_\mu$$

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NEUTRINOS REMAIN MOST ELUSIVE

Fundamental questions remain:

- Are neutrinos Dirac/Majorana? Leptogenesis?
- The three-mass ordering?
- How large is the CP phase?
- The next discovery (maybe)
- Are there “Non-Standard Interactions” (NSI)?

1. NSI at the COHERENT & LHC #
2. NSI with a “Leptonic Scalar” *

arXiv:1910.03272 [hep-ph] JHEP 1911 (2019), 028
TH, Hongkai Liu, Jiajun Liao, Danny Marfatia

* arXiv:1910.01132 [hep-ph]
Andre de Gouvea, Bhupal Dev, Bhaska Dutta, T.
Ghosh, TH, Yongchao Zhang

1. “NSI”: COHERENT & LHC

As originally formulated by Wolfenstein, NSI:

$$\begin{aligned}\mathcal{L}_{\text{NSI}} &= -2\sqrt{2}G_F \sum_{f,C,\alpha,\beta} \epsilon_{\alpha\beta}^{fP} (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{f} \gamma_\mu P_C f) \\ &= -\sqrt{2}G_F \epsilon_{\alpha\beta}^{fV} (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{f} \gamma_\mu f) - \sqrt{2}G_F \epsilon_{\alpha\beta}^{fA} (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{f} \gamma_\mu \gamma^5 f) \\ \epsilon_{\alpha\beta}^{fV} &\equiv \epsilon_{\alpha\beta}^{fL} + \epsilon_{\alpha\beta}^{fR}, \quad \epsilon_{\alpha\beta}^{fA} \equiv \epsilon_{\alpha\beta}^{fR} - \epsilon_{\alpha\beta}^{fL}\end{aligned}$$

- We will only consider the NC NSI.
- For a heavy mediator: $\epsilon \sim g'^2 v_{\text{EW}}^2/M^2$.
- Interesting to consider UV formulation

L. Wolfenstein (1978);
T. Ohlsson arXiv:1209.2710;
Farzan & Tortola, arXiv:1710.09160.

Consider a UV-complete $U(1)'$ model:

$$\mathcal{L} = \mathcal{L}_{\text{SM}} - \frac{1}{4} Z'^{\mu\nu} Z'_{\mu\nu} + \frac{1}{2} M_{Z'}^2 Z'^\mu Z'_\mu + Z'_\mu J_X^\mu$$
$$J_X^\mu = g' \left[\sum_q Q'_q \bar{q} \gamma^\mu q + \sum_{L_\ell = \nu_{\ell L}, \ell} Q'_\ell \bar{L}_\ell \gamma^\mu L_\ell \right]$$

the quark charges $Q'_{1,2,3}$ and lepton charges $Q'_{e,\mu,\tau}$ satisfying the constraint

$$3(Q'_1 + Q'_2 + Q'_3) + Q'_e + Q'_\mu + Q'_\tau = 0.$$
$$Q'_1 = Q'_2 = Q'_3 = Q'_q$$

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Our representative choices for $U(1)'$ charges:

(A) $Q'_q = 1/3, \underline{Q'_\mu = -3, Q'_e = Q'_\tau = 0}.$

(B) $Q'_q = 1/3, \underline{Q'_\mu = Q'_\tau = -3/2, Q'_e = 0}.$

(C) $Q'_q = 1/3, \underline{Q'_\tau = -3, Q'_e = Q'_\mu = 0}.$

- Electron flavor heavily constrained (beam-dump), not included here
- $M_{Z'} \sim 5 \text{ MeV} - \mathcal{O}(\text{TeV})$
→ heavy for oscillation expts, but suitable for LHC searches.

Neutrino oscillation experiments

$$H = \frac{1}{2E} U \begin{pmatrix} 0 & 0 & 0 \\ 0 & \delta m_{21}^2 & 0 \\ 0 & 0 & \delta m_{31}^2 \end{pmatrix} U^\dagger + V, \quad V = \sqrt{2} G_F N_e \begin{pmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{e\tau}^* & \epsilon_{\mu\tau}^* & \epsilon_{\tau\tau} \end{pmatrix}$$

$\epsilon_{\alpha\alpha}^u$	Current data	DUNE+T2HK
ϵ_{ee}^u	$[-1.192, -0.802] \oplus [-0.020, +0.456]$	$[-0.407, -0.270] \oplus [-0.072, +0.064]$
$\epsilon_{\mu\mu}^u$	$[-0.130, 0.152]$	$[-0.019, +0.018]$
$\epsilon_{\tau\tau}^u$	$[-0.152, 0.130]$	$[-0.017, +0.017]$

I. Esteban et al., arXiv:1805.04530.

Table 1: 2σ allowed ranges for the diagonal NSI parameters from the global analysis of current oscillation data assuming both LMA and LMA-D [22] and from the simulation of next generation neutrino oscillation experiments DUNE and T2HK.

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In our Z' model: $M_{Z'} \sim 5 \text{ MeV} - 20 \text{ MeV}$

$$\mathcal{L}_{\text{eff}} = \frac{(g')^2}{M_{Z'}^2} \left[\sum_q Q'_q \bar{q} \gamma^\mu q \right] \left[\sum_\alpha Q'_\alpha \bar{\nu}_\alpha \gamma^\mu P_L \nu_\alpha \right] \quad \epsilon_{\alpha\alpha}^{qV} = -\frac{(g')^2 Q'_\alpha Q'_q}{\sqrt{2} G_F M_{Z'}^2}$$

$$\epsilon_{\alpha\beta} \equiv \sum_q \epsilon_{\alpha\beta}^{qV} \frac{N_q}{N_e}$$

CEvNS Constraints – COHERENT expt

$$M_{Z'} \sim 10 \text{ MeV} - 10 \text{ GeV}$$

detector

PDF

SM

$$\pi^+ \rightarrow \mu^+ + \nu_\mu, \text{ and}$$

$$\mu^+ \rightarrow e^+ + \bar{\nu}_\mu + \nu_e.$$

$$N_{th}(t, E_r, \epsilon) = \sum_{\alpha} \frac{m_{\text{det}} N_A}{M} \int_{\Delta E_r} dE_r \int_{\Delta t} dt \rho_{\alpha}(t) \int_{E_{\nu}^{\min}}^{E_{\nu}^{\max}} dE_{\nu} \phi_{\alpha}(E_{\nu}) \frac{d\sigma_{\alpha}(\epsilon)}{dE_r}$$

$$\frac{d\sigma_{\alpha}(\epsilon)}{dE_r} = \frac{G_F^2}{2\pi} Q_{\alpha}^2 F^2(Q^2) M \left(2 - \frac{ME_r}{E_{\nu}^2}\right),$$

$$Q_{\alpha}^2 = [Z(g_p^V + 2\epsilon_{\alpha\alpha}^{uV} + \epsilon_{\alpha\alpha}^{dV}) + N(g_n^V + \epsilon_{\alpha\alpha}^{uV} + 2\epsilon_{\alpha\alpha}^{dV})]^2$$

$$\phi_{\nu_{\mu}}(E_{\nu_{\mu}}) = \frac{2m_{\pi}}{m_{\pi}^2 - m_{\mu}^2} \delta\left(1 - \frac{2E_{\nu_{\mu}}m_{\pi}}{m_{\pi}^2 - m_{\mu}^2}\right)$$

$$\phi_{\nu_e}(E_{\nu_e}) = \frac{192}{m_{\mu}} \left(\frac{E_{\nu_e}}{m_{\mu}}\right)^2 \left(\frac{1}{2} - \frac{E_{\nu_e}}{m_{\mu}}\right),$$

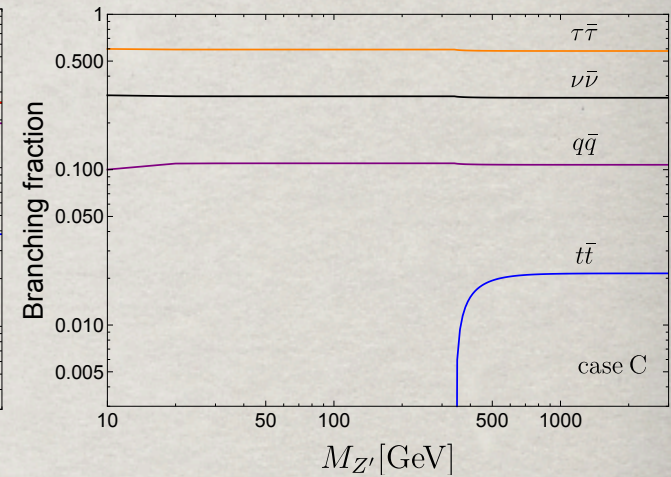
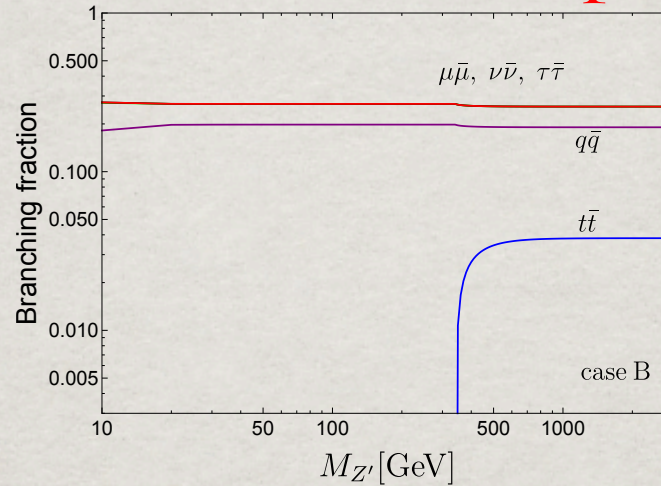
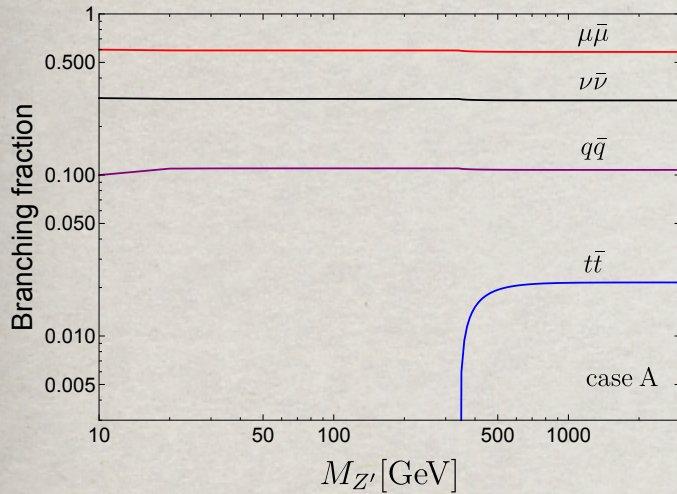
$$\phi_{\bar{\nu}_{\mu}}(E_{\bar{\nu}_{\mu}}) = \frac{64}{m_{\mu}} \left(\frac{E_{\bar{\nu}_{\mu}}}{m_{\mu}}\right)^2 \left(\frac{3}{4} - \frac{E_{\bar{\nu}_{\mu}}}{m_{\mu}}\right),$$

$$\frac{d\sigma_{\alpha, \text{CsI}}}{dE_r} = \frac{d\sigma_{\alpha, \text{Cs}}}{dE_r} + \frac{d\sigma_{\alpha, \text{I}}}{dE_r} \quad \epsilon_{ee}^{uV} = \epsilon_{ee}^{dV} = \frac{g'^2 Q'_q Q'_e}{\sqrt{2} G_F (2ME_r + M_{Z'}^2)}$$

$$\epsilon_{\mu\mu}^{uV} = \epsilon_{\mu\mu}^{dV} = \frac{g'^2 Q'_q Q'_\mu}{\sqrt{2} G_F (2ME_r + M_{Z'}^2)}$$

$$\chi^2 = \sum_{i=4}^{15} \left[\frac{N_{\text{meas}}^i - N_{\text{th}}^i(1 + \gamma) - B_{\text{on}}(1 + \beta)}{\sigma_{\text{stat}}^i} \right]^2 + \left(\frac{\gamma}{\sigma_{\gamma}} \right)^2 + \left(\frac{\beta}{\sigma_{\beta}} \right)^2$$

LHC Studies: $Z' \rightarrow l^+ l^-$ (Cases A,B,C on p.5)



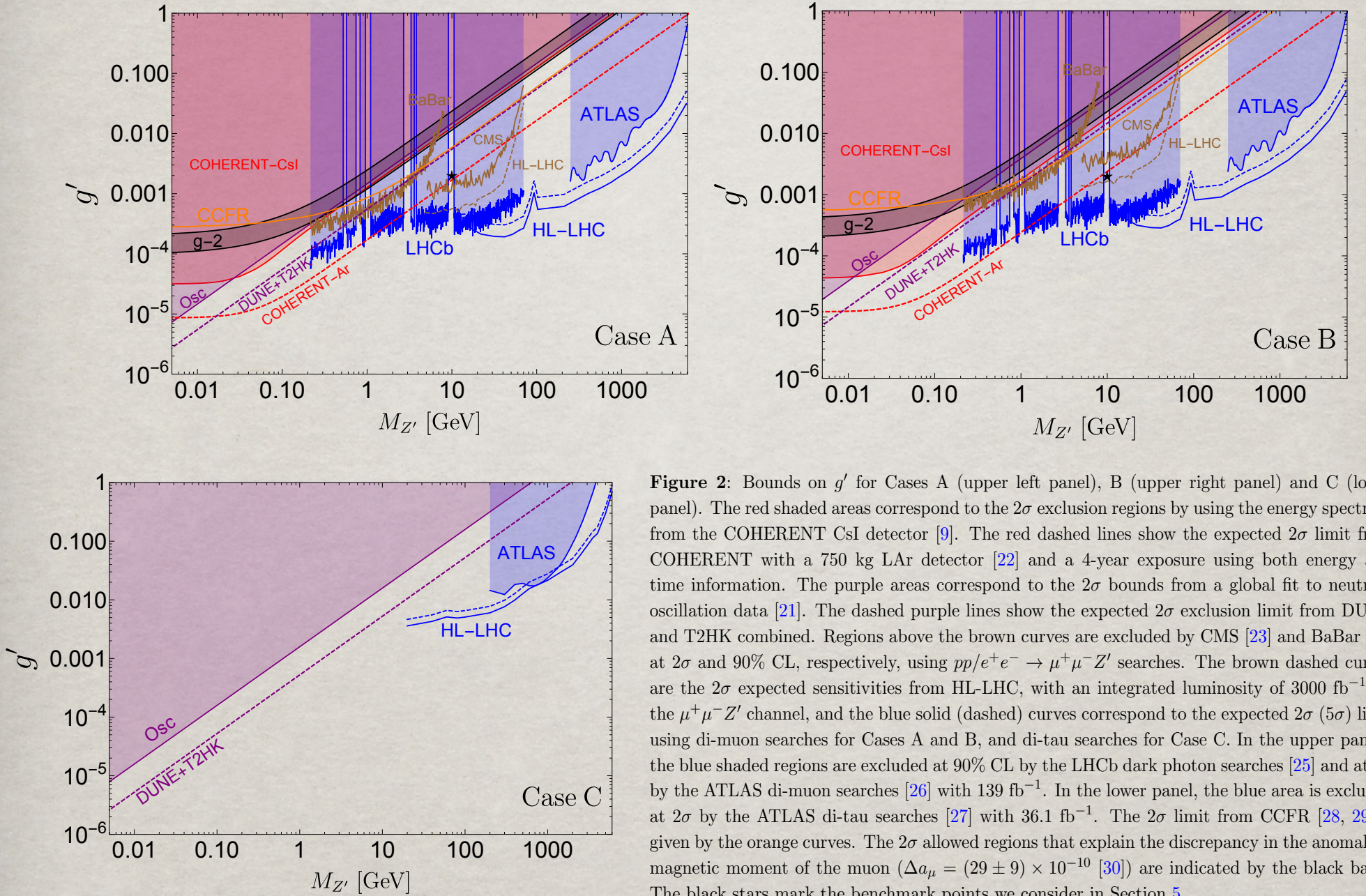
For $M_{Z'} < M_Z$ $pp \rightarrow Z^*/\gamma^* \rightarrow l^+ l^- + Z' \rightarrow l^+ l^- + l^+ l^- + X$.

- LHCb & CMS 4-lepton recast

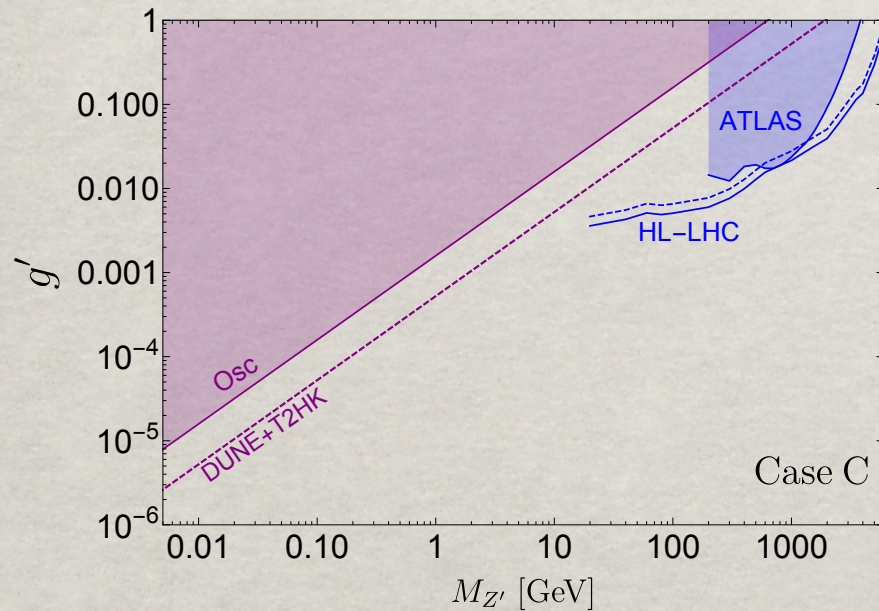
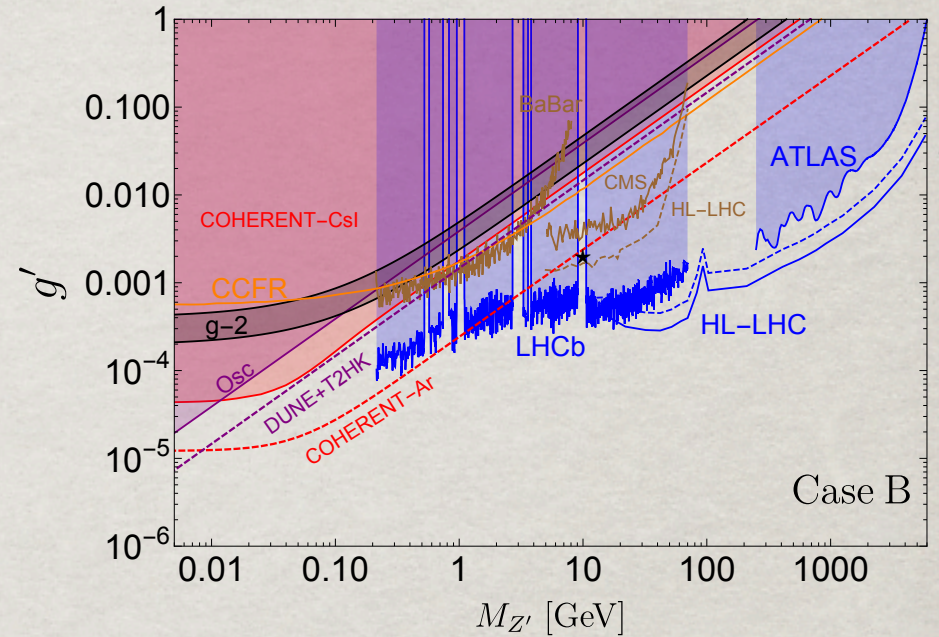
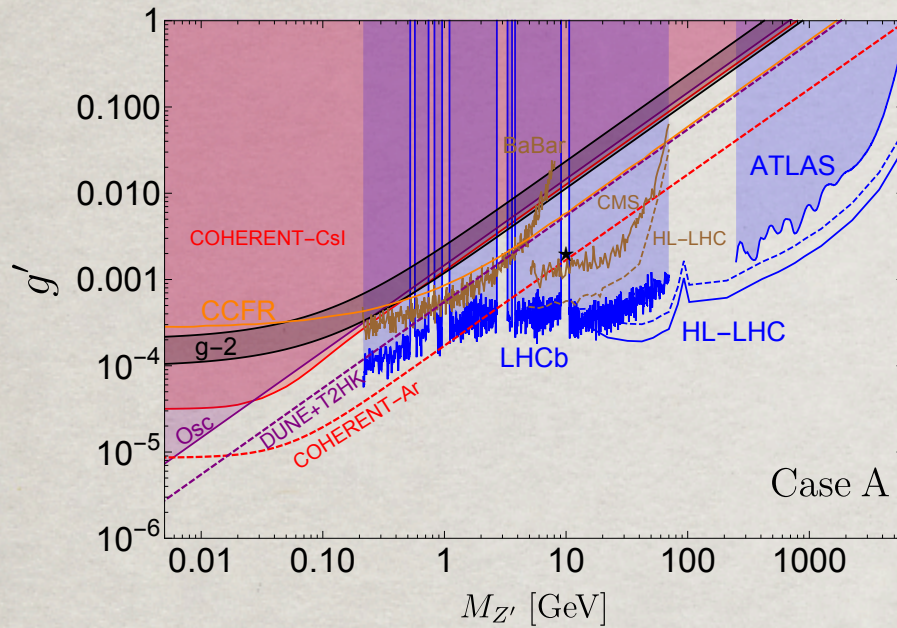
For $M_{Z'} > M_Z$ $pp \rightarrow Z' \rightarrow l^+ l^- + X$

- ATLAS/CMS existing results recast
- HL-LHC new study

Everything together



Everything together

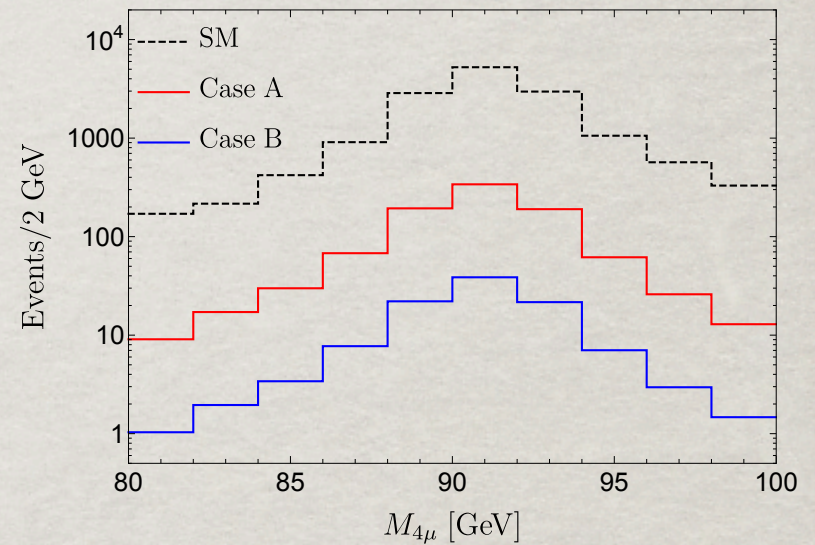
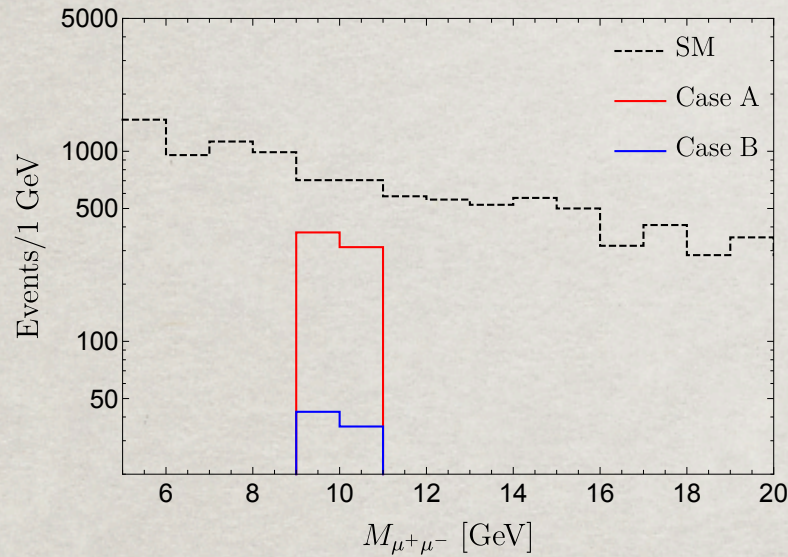


- $M_{Z'} \sim 5 \text{ MeV} - 20 \text{ MeV}$:
DUNE+T2HK
- $M_{Z'} \sim 20 \text{ MeV} - 1 \text{ GeV}$:
COHERENT Lar
- $M_{Z'} \sim 1 \text{ GeV} - 20 \text{ GeV}$:
LHCb, CMS 4-leptons
- $M_{Z'} \sim 20 \text{ GeV} - 4 \text{ TeV}$:
HL-LHC 2-leptons

COHERENT & LHC: Correlated Studies

- Generate events for SM+Z'
- Scan & search for the signal

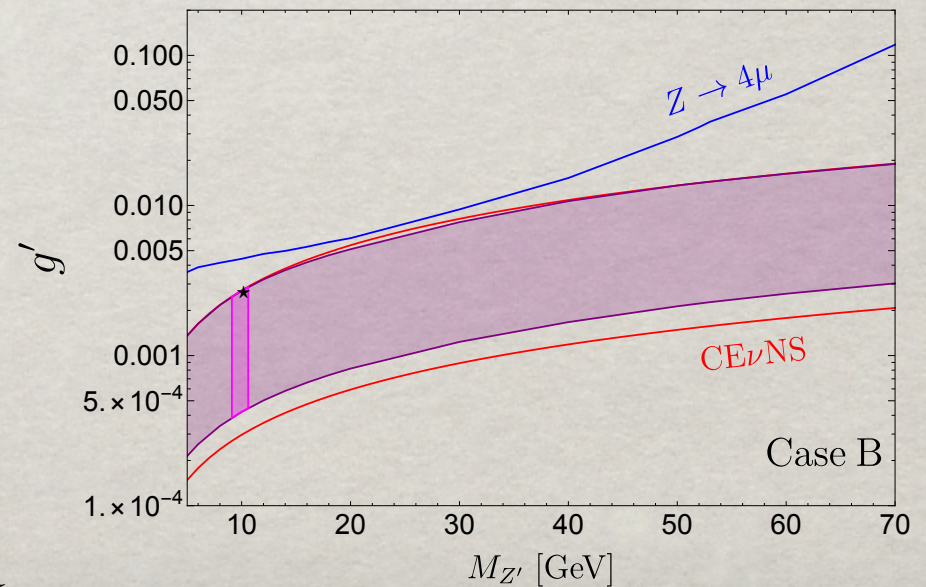
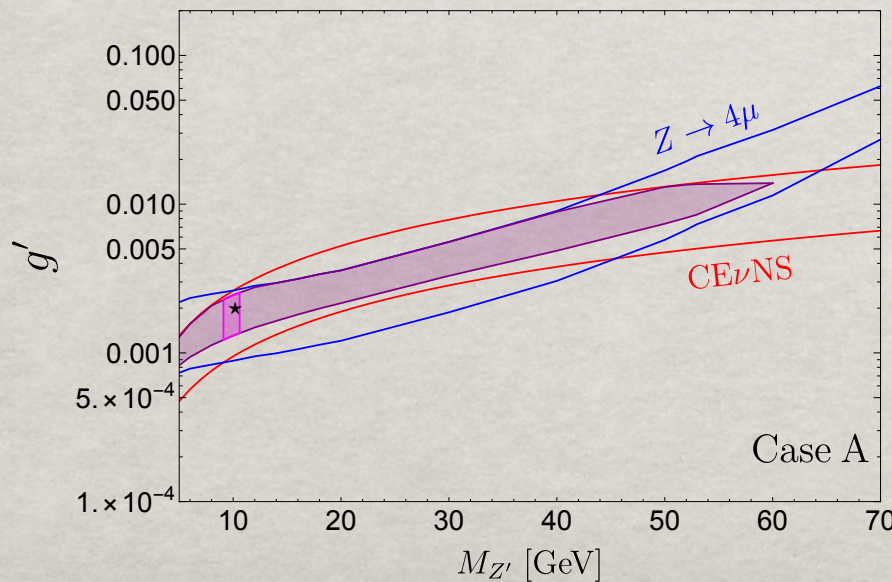
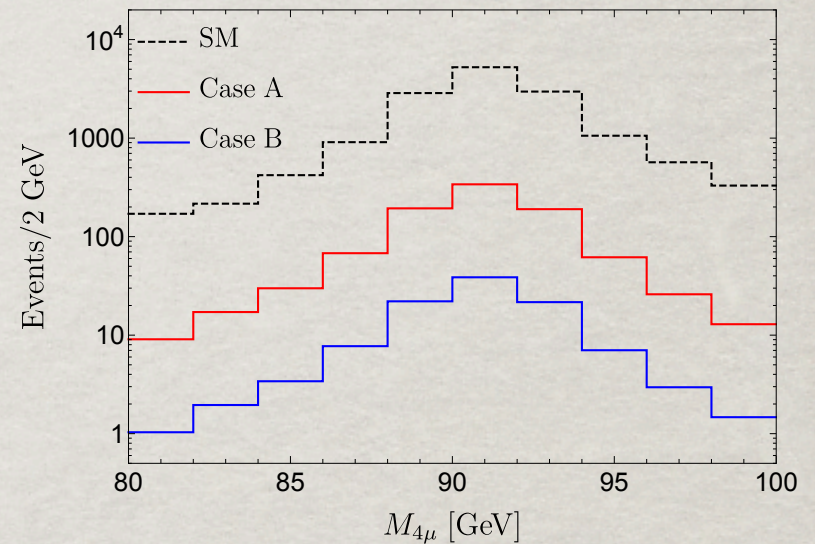
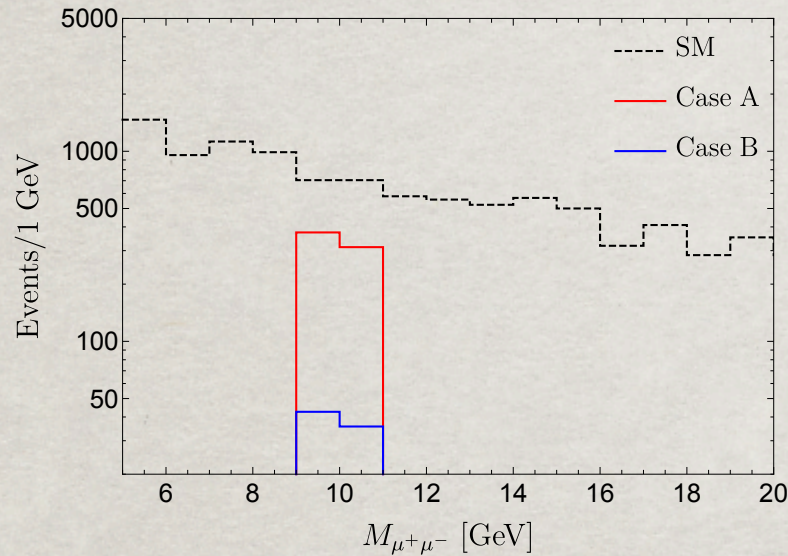
$$\chi^2 = \sum_i \frac{N_{S,i}^2}{N_{B,i} + (\sigma_B N_{B,i})^2}$$



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2. NSI with a “Leptonic Scalar”

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Tree-level “seesaw mechanism”

The dim-5 Weinberg operator:

$$\frac{1}{\Lambda} (y_\nu LH)(y_\nu LH) + h.c. \Rightarrow \frac{y_\nu^2 v^2}{\Lambda} \bar{\nu}_L \nu_R^c.$$



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(along with \mathbf{W}_R)



Minkowski (1976); Yanagita (1979); Glashow (1980); Mohapatra, Senjanovic (1980);
Magg, Wetterich (1980); Lasarides, Shafi (1981); Mohapatra, Senjanovic (1981);
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Type-II: Add a scalar triplet $\Phi (Y = 2) : \phi^{\pm\pm}, \phi^\pm, \phi^0$ (along with \mathbf{W}_R)

$$Y_{ij} L_i^T C(i\sigma_2) \Phi L_j + h.c. \quad \mu H^T (i\sigma_2) \Phi^\dagger H + h.c.$$



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 $Y_{ij} L_i^T C(i\sigma_2) \Phi L_j + h.c. \quad \mu H^T (i\sigma_2) \Phi^\dagger H + h.c.$

Type-III: Add a fermionic triplet $T (Y = 0) : T^+ T^0 T^-$
 $-M_T(T^+ T^- + T^0 T^0/2) + y_T^i H^T i\sigma_2 T L_i + h.c.$

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2. NSI with a “Leptonic Scalar”

Neutrinos are elusive, and could couple to new particles, and thus modify their behaviors. Consider:

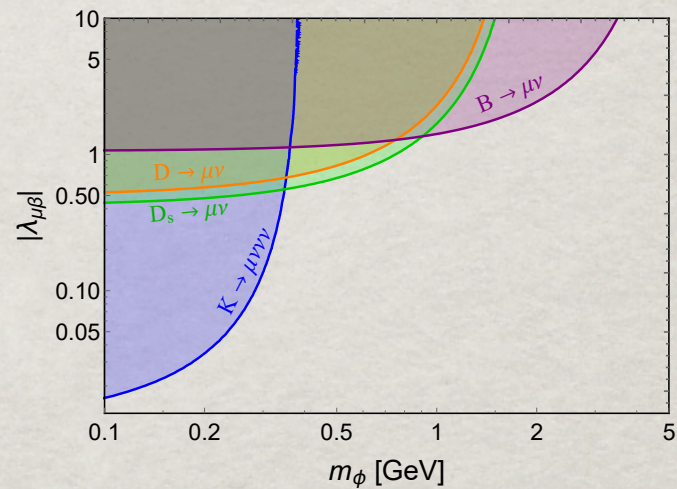
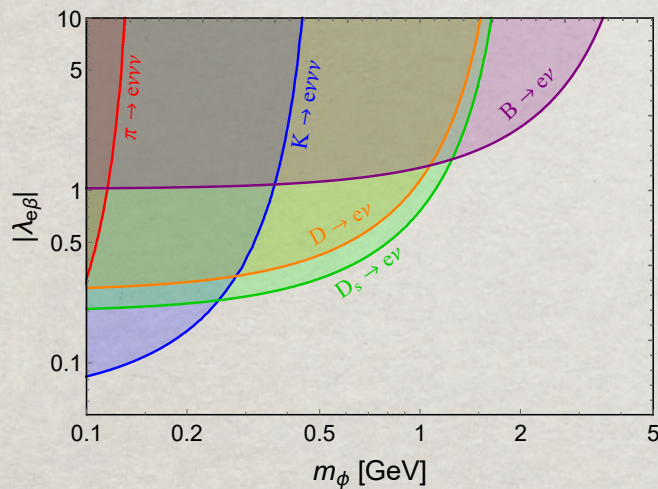
$$\mathcal{L} \supset \frac{1}{2} \lambda_{\alpha\beta} \phi \nu_{\alpha} \nu_{\beta}$$

- ϕ carries lepton-number $L = -2$
- At renormalizable level: $\phi \nu_R \nu_R$
- At dim-6: $\lambda_{\alpha\beta} \sim \kappa_1 \kappa_2 v_{EW}^2/M^2$
- Could be from a UV complete formulation
- It can radiate off any neutrino and thus could effect many processes:
 - meson decays; W/Z decays
 - light DM searches; IceCube
 - collider experiments

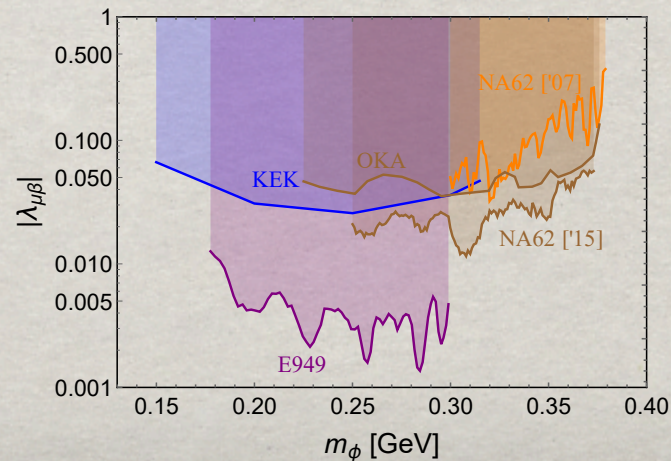
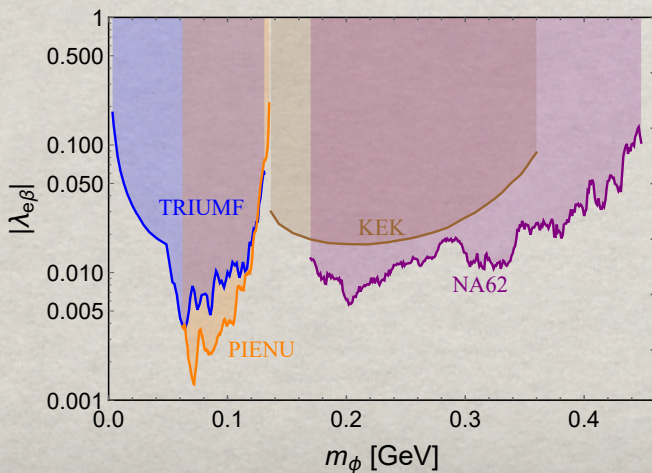
Low energy bounds: Meson decays

For leptonic decays of charged mesons $P^- \rightarrow \ell^- \bar{\nu}$ with $P^- = \pi^-, K^-, D^-, D_S^-, B^-$,

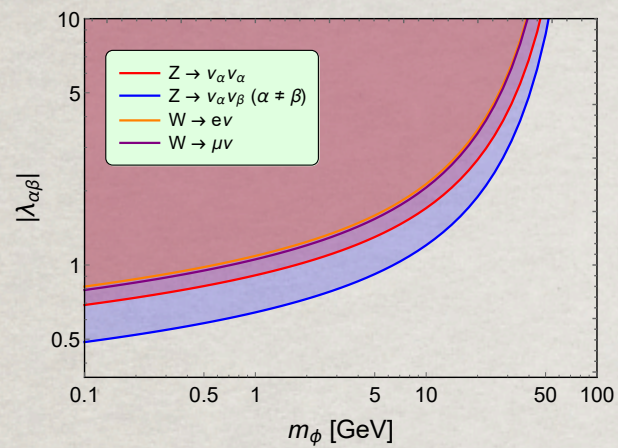
$$\Gamma(P^- \rightarrow \ell^- \bar{\nu} \phi) = \frac{G_F^2 |V_{qq'}|^2 m_P^3 f_P^2 \sum_{\beta} |\lambda_{\alpha\beta}|^2}{256\pi^3} \times \int_{x_{\phi}}^{(1-\sqrt{x_{\ell}})^2} dx \frac{((x+x_{\ell}) - (x-x_{\ell})^2)(x-x_{\phi})^2}{x^3} \lambda^{1/2}(1, x, x_{\ell})$$



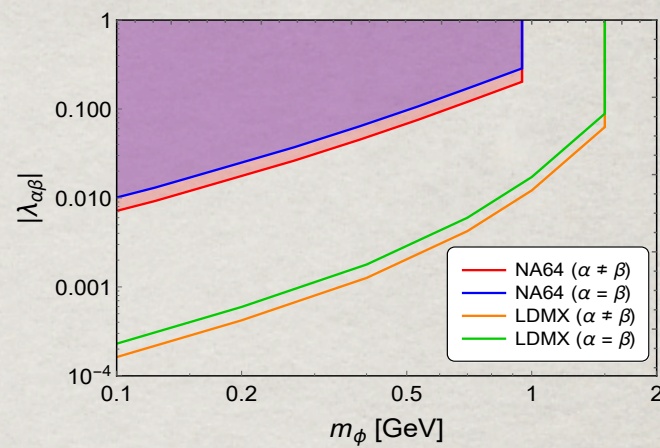
Heavy neutrino searches in meson decay spectra



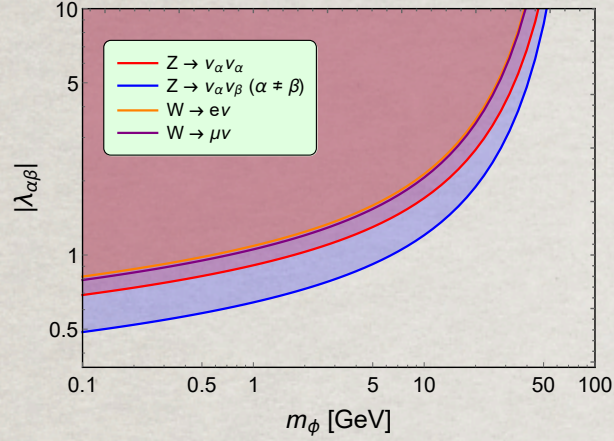
W/Z decays:



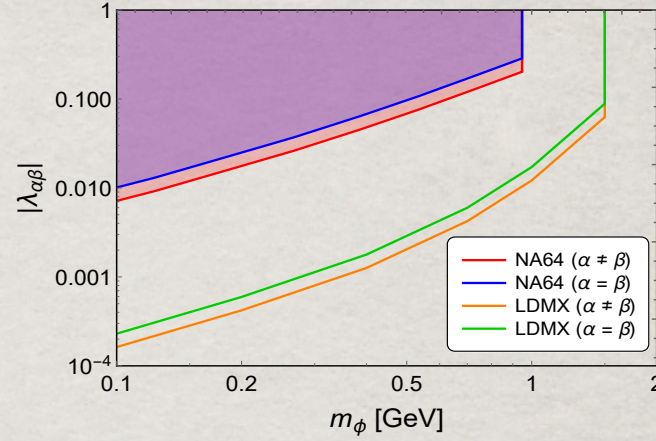
Dark photon/DM searches:



W/Z decays:



Dark photon/DM searches:



Ref.	Process	Data	Couplings	Mass range
[1, 2]	$\pi^- \rightarrow e^- \bar{\nu}_e \nu \bar{\nu}$	$\text{BR} < 5 \times 10^{-6}$	$\sum_{\beta} \lambda_{e\beta} ^2$	$m_{\phi} < 131 \text{ MeV}$
[1, 2]	$K^- \rightarrow e^- \bar{\nu}_e \nu \bar{\nu}$	$\text{BR} < 6 \times 10^{-5}$	$\sum_{\beta} \lambda_{e\beta} ^2$	$m_{\phi} < 444 \text{ MeV}$
[1, 2]	$K^- \rightarrow \mu^- \bar{\nu}_{\mu} \nu \bar{\nu}$	$\text{BR} < 2.4 \times 10^{-6}$	$\sum_{\beta} \lambda_{\mu\beta} ^2$	$m_{\phi} < 386 \text{ MeV}$
[1, 2]	$D^- \rightarrow e^- \bar{\nu}_e$	$\text{BR} < 8.8 \times 10^{-6}$	$\sum_{\beta} \lambda_{e\beta} ^2$	$m_{\phi} < 1.52 \text{ GeV}$
[1, 2]	$D^- \rightarrow \mu^- \bar{\nu}_{\mu}$	$\text{BR} < 3.4 \times 10^{-5}$	$\sum_{\beta} \lambda_{\mu\beta} ^2$	$m_{\phi} < 1.39 \text{ GeV}$
[1, 21]	$D_s^- \rightarrow e^- \bar{\nu}_e$	$\text{BR} < 8.3 \times 10^{-5}$	$\sum_{\beta} \lambda_{e\beta} ^2$	$m_{\phi} < 1.64 \text{ GeV}$
[1, 21]	$D_s^- \rightarrow \mu^- \bar{\nu}_{\mu}$	$\text{BR} = (5.50 \pm 0.23) \times 10^{-3}$	$\sum_{\beta} \lambda_{\mu\beta} ^2$	$m_{\phi} < 1.50 \text{ GeV}$
[1, 21]	$B^- \rightarrow e^- \bar{\nu}_e$	$\text{BR} < 9.8 \times 10^{-7}$	$\sum_{\beta} \lambda_{e\beta} ^2$	$m_{\phi} < 3.54 \text{ GeV}$
[1, 21]	$B^- \rightarrow \mu^- \bar{\nu}_{\mu}$	$\text{BR} = (2.90 - 10.7) \times 10^{-7}$	$\sum_{\beta} \lambda_{\mu\beta} ^2$	$m_{\phi} < 3.50 \text{ GeV}$
[1, 20]	$\tau^- \rightarrow e^- \bar{\nu}_e \nu_{\tau}$	$\text{BR} = (17.82 \pm 0.04)\%$	$\sum_{\beta} \lambda_{e\beta} ^2$	$m_{\phi} < 741 \text{ MeV}$
[1, 20]	$\tau^- \rightarrow \mu^- \bar{\nu}_{\mu} \nu_{\tau}$	$\text{BR} = (17.39 \pm 0.04)\%$	$\sum_{\beta} \lambda_{\mu\beta} ^2$	$m_{\phi} < 741 \text{ MeV}$
[1, 21]	$P^- \rightarrow e^- N$	see Ref. [25]	$\sum_{\beta} \lambda_{e\beta} ^2$	$3.3 \text{ MeV} < m_{\phi} < 448 \text{ MeV}$
[1, 21]	$P^- \rightarrow \mu^- N$	see Ref. [25]	$\sum_{\beta} \lambda_{\mu\beta} ^2$	$87 \text{ MeV} < m_{\phi} < 379 \text{ MeV}$
[1]	$Z \rightarrow \text{inv.}$	$\text{BR} = (20.0 \pm 0.055)\%$	$\sum_{\alpha, \beta} S_{\alpha\beta} \lambda_{\alpha\beta} ^2$	$m_{\phi} < 52.2 \text{ GeV}$
[1]	$W \rightarrow e \nu$	$\text{BR} = (10.71 \pm 0.16)\%$	$\sum_{\beta} \lambda_{e\beta} ^2$	$m_{\phi} < 38.8 \text{ GeV}$
[1]	$W \rightarrow \mu \nu$	$\text{BR} = (10.63 \pm 0.15)\%$	$\sum_{\beta} \lambda_{\mu\beta} ^2$	$m_{\phi} < 39.3 \text{ GeV}$
[2]	MINOS	see Ref. [2]	$ \lambda_{\mu\mu} $	$m_{\phi} < 1.67 \text{ GeV}$
[2]	DUNE	see Ref. [2]	$ \lambda_{\mu\mu} $	$m_{\phi} < 3.00 \text{ GeV}$
[26]	NA64	see Ref. [26]	$\sum_{\alpha, \beta} S_{\alpha\beta} \lambda_{\alpha\beta} ^2$	$m_{\phi} < 948 \text{ MeV}$
[27]	LDMX	see Ref. [27]	$\sum_{\alpha, \beta} S_{\alpha\beta} \lambda_{\alpha\beta} ^2$	$m_{\phi} < 1.50 \text{ GeV}$
[28, 29]	IceCube	see Ref. [28]	$ \lambda_{\alpha\beta} $	$m_{\phi} < 2.0 (15.0) \text{ GeV}$

Some other constraints:

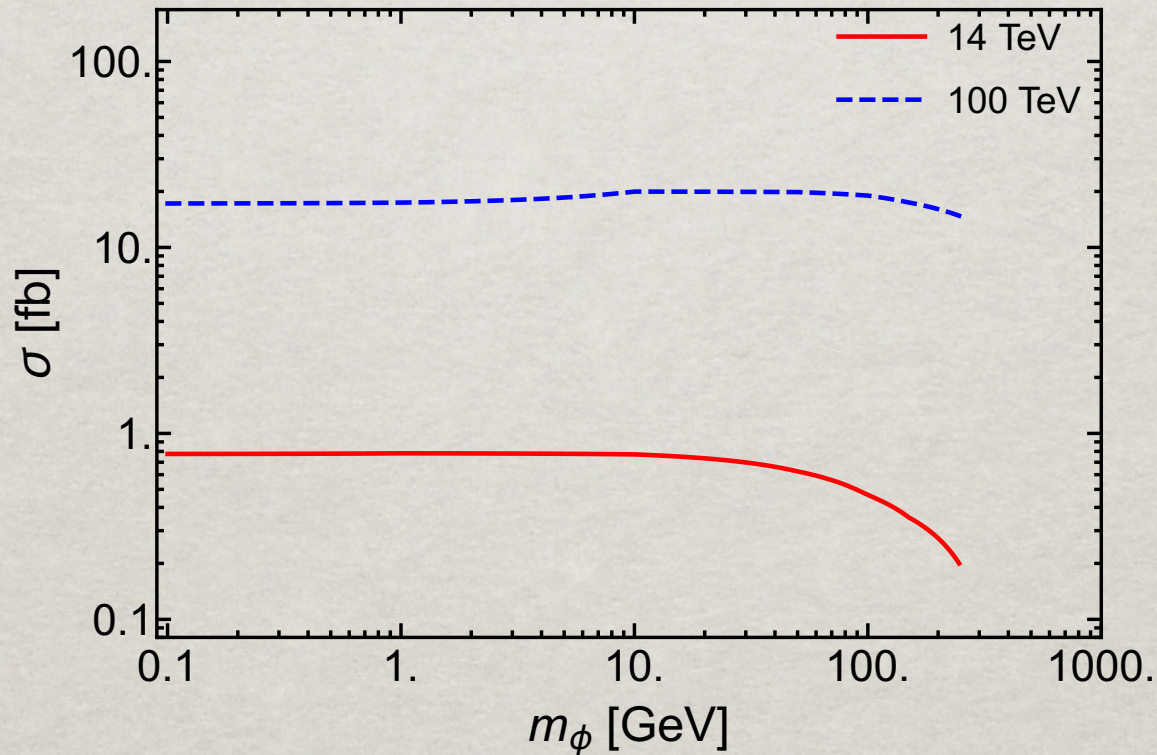
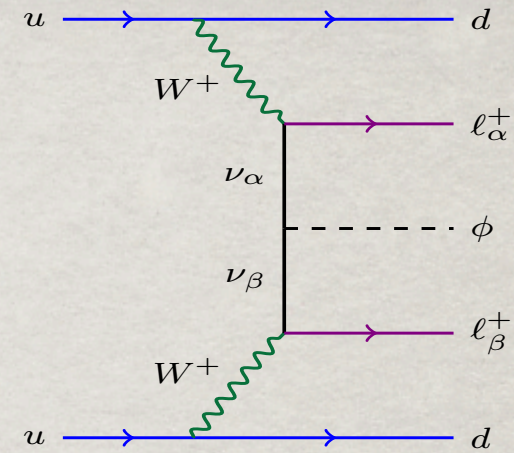
- *Muon decay*: As discussed in Section 2.4, ϕ can be emitted from tree-level decay $\mu \rightarrow e\nu\nu\phi$. As a result of the precise μ decay data, for sufficiently light ϕ , the limits from μ decay are expected to be much more stringent than those from τ decays. In addition, the electron [98, 99] and neutrino [100, 101] spectra could be altered in presence of ϕ , which can also be used to set limits on $\lambda_{\alpha\beta}$.
- *Tritium decay*: If the scalar mass $m_\phi \lesssim \mathcal{O}(10 \text{ eV})$, it can be produced from tritium decay in the process ${}^3\text{H} \rightarrow {}^3\text{He}^+ + e^- + \nu + \phi$ [102], and this process can be probed in the KATRIN experiment [103, 104].
- *$0\nu\beta\beta$ decay*: The coupling of ϕ to electron neutrinos contributes to $0\nu\beta\beta$ decays via the process $(Z, A) \rightarrow (Z + 2, A)e^-e^-\phi$ if the mass $m_\phi \lesssim \mathcal{O}(\text{MeV})$ – the typical Q -value for the relevant nuclei. This is strongly constrained by the searches of Majoron emission in $0\nu\beta\beta$ decay experiments like NEMO-3 using ${}^{100}\text{Mo}$ [6, 7, 11] and ${}^{150}\text{Nd}$ [9] nuclei, as well as KamLAND-Zen [12] and EXO-200 [13] using ${}^{136}\text{Xe}$. Somewhat weaker limits were also obtained by NEMO-3 using ${}^{48}\text{Ca}$ [8] and ${}^{82}\text{Se}$ [10], as well as by GERDA using ${}^{76}\text{Ge}$ [14].
- *Supernovae*: A light ϕ can be produced abundantly in the supernova core if its mass $m_\phi \lesssim \mathcal{O}(30 \text{ MeV})$ – the typical core temperature of supernovae. The couplings $|\lambda_{\alpha\beta}|$ can be constrained from both the luminosity and deleptonization arguments [105–107].
- *CMB and BBN*: As a light particle, ϕ itself contributes to the relativistic degrees of freedom N_{eff} if the mass $m_\phi \lesssim 100 \text{ keV}$ [108]. The current precision cosmological data $\Delta N_{\text{eff}} = 0.18$ at 1σ C.L. [109] has excluded a large parameter space for such light leptonic scalar mass m_ϕ and the couplings $|\lambda_{\alpha\beta}|$. Similarly, the big-bang-nucleosynthesis (BBN) constraints rule out $m_\phi \lesssim 0.2 \text{ MeV}$ for sizable couplings $\lambda_{\alpha\beta}$, as long as they allow ϕ particles to thermalize at BBN temperature [110].
- *Neutrino decay*: For sufficiently light ϕ , the heavier neutrinos might decay via $\nu_j \rightarrow \nu_i + \phi$ with the mass indices $i, j = 1, 2, 3$ and $i < j$. Therefore we can impose stringent bounds on the leptonic scalar mass m_ϕ and the λ_{ij} couplings from the solar neutrino data [111–115]. There are also constraints from atmospheric and long baseline experiments [116–118]. The CMB limits on neutrino free streaming could also set limits on neutrino decays, as long as the mediator is lighter than neutrino mass and the non-diagonal couplings λ_{ij} are non-vanishing [93, 94, 119].

LHC searches:

A unique, clean channel

$$pp \rightarrow \ell_{\alpha}^{\pm} \ell_{\beta}^{\pm} \phi jj$$

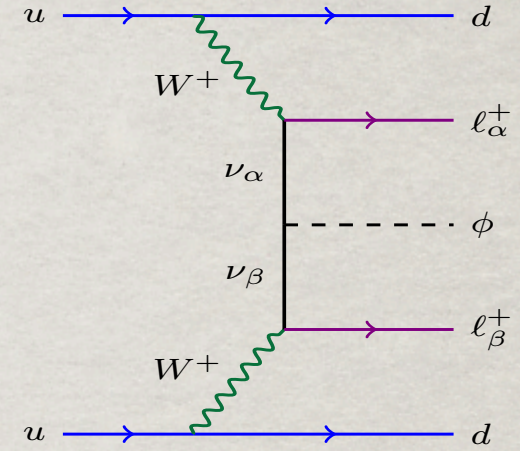
$$\lambda_{\alpha\beta}=1 \ (\alpha,\beta=e,\mu)$$



(m_{ϕ} not to exceed the EW scale)

Signal & Backgrounds:

- the EW process $pp \rightarrow W^\pm W^\pm jj \rightarrow jj \ell_\alpha^\pm \ell_\beta^\pm \nu\nu$,
- the QCD process $pp \rightarrow W^\pm W^\pm jj \rightarrow jj \ell_\alpha^\pm \ell_\beta^\pm \nu\nu$
- $pp \rightarrow W^\pm Z jj \rightarrow jj \ell_\alpha^\pm \ell_\beta^\pm \ell_\beta^\mp \nu$,



Channels		$e^\pm e^\pm$	$e^\pm \mu^\pm$	$\mu^\pm \mu^\pm$	Total
Signal		40	129	84	253
$W^\pm W^\pm jj$ (EW)		37	137	89	263
$W^\pm W^\pm jj$ (QCD)		2	9	2	13
$W^\pm Z jj$		29	94	54	177
Total background		68	240	145	453
Significance	syst. error 0%	3.87	6.73	5.53	9.53
	syst. error 10%	3.24	4.21	4.00	4.83

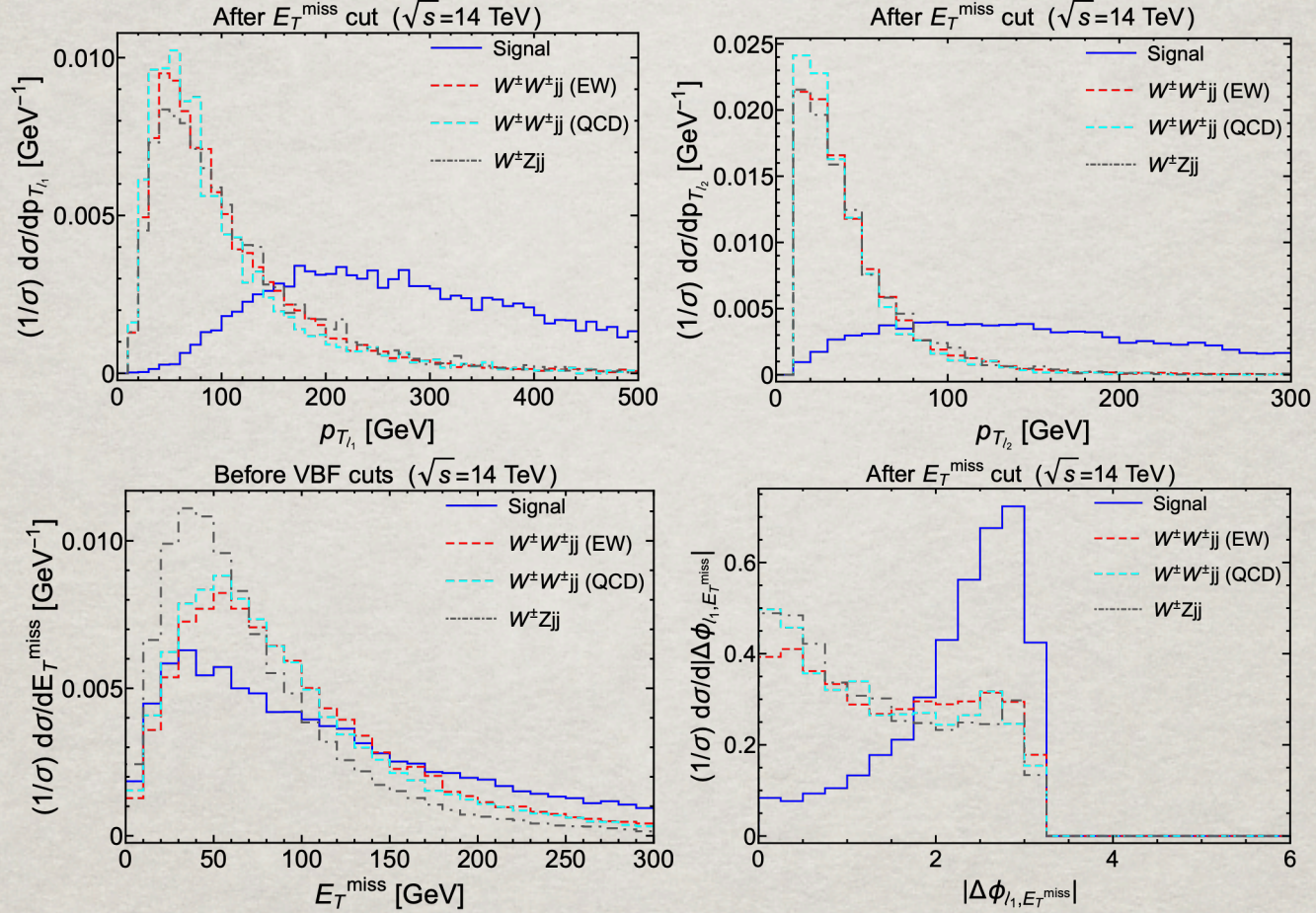


Figure 8. Kinematic distributions of the signal and SM backgrounds $W^\pm W^\pm jj$ (EW), $W^\pm W^\pm jj$ (QCD) and $W^\pm Zjj$ after the E_T^{miss} cut. The top left and right panels are respectively for the p_T distributions of the leading lepton ℓ_1 and the sub-leading lepton ℓ_2 , and the lower left and right panels respectively for the missing transverse energy E_T^{miss} and the angular separation $|\Delta\phi_{\ell_1 E_T^{\text{miss}}}|$. Only the E_T^{miss} distribution is shown before VBF cuts.

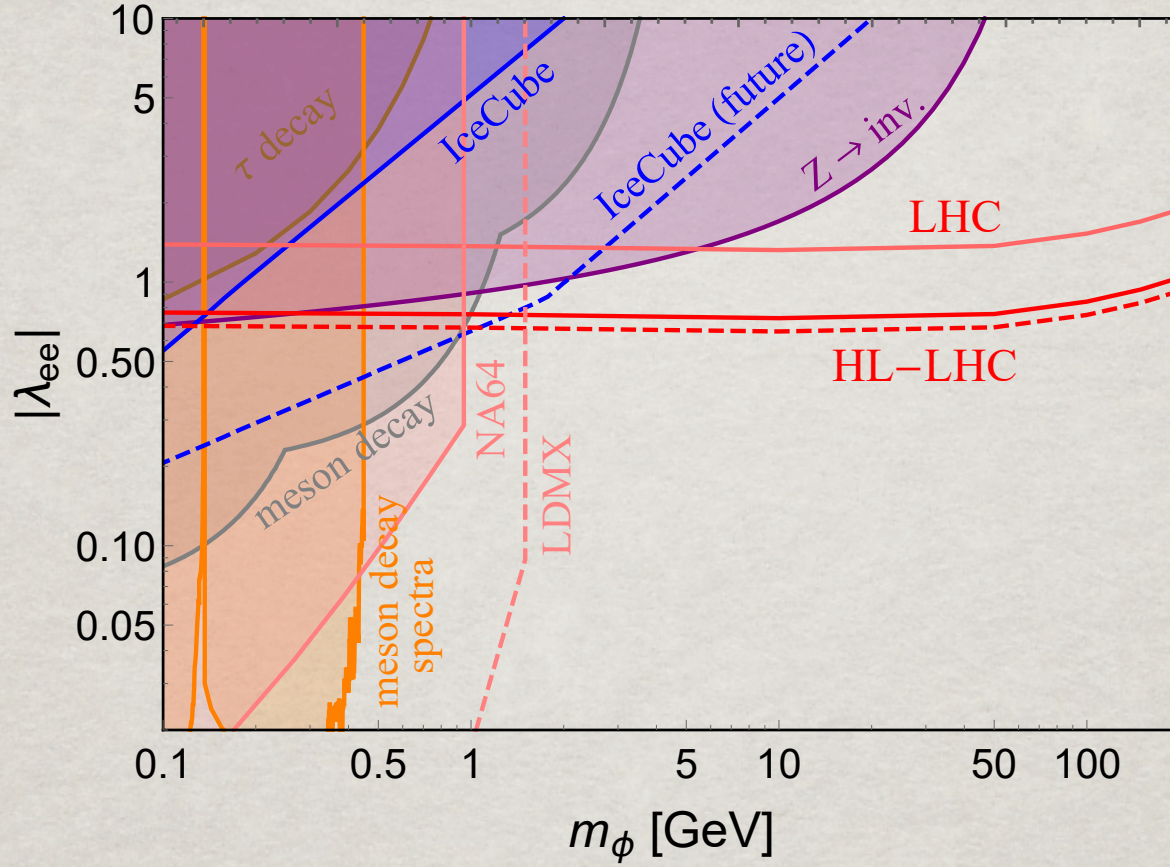
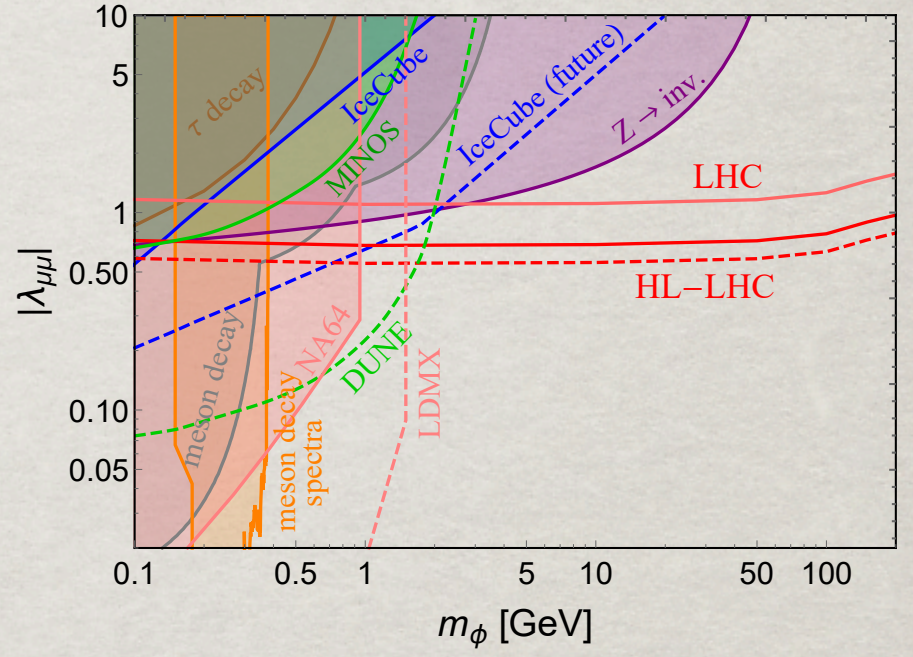
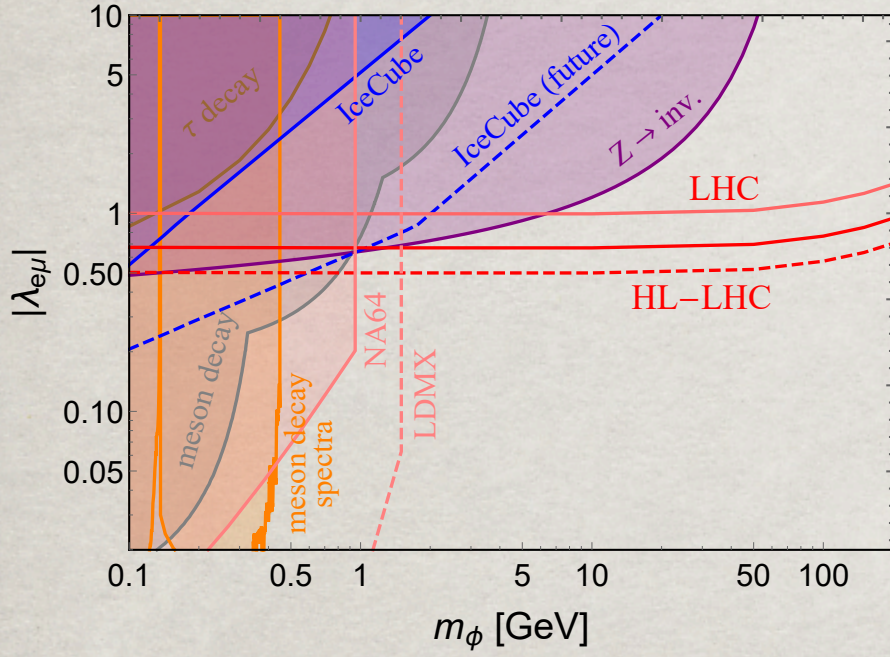


Figure 9. Prospects of the coupling $|\lambda_{ee}|$ as a function of the scalar mass m_ϕ at 14 TeV LHC with luminosity of 300 fb^{-1} (solid thin red line) and HL-LHC with 3 ab^{-1} and with systematic errors of 10% (solid thick red line) and 0% (dashed thick red line). Also shown are the low-energy limits (cf. Table 1) from meson decay (gray), τ decay (brown), heavy neutrino searches in meson decay spectra (orange), invisible Z decay (purple), light DM searches in NA64 (pink) and the prospects at LDMX (dashed pink), the current IceCube limits on neutrino–neutrino interactions (blue) and prospects (dashed blue). All the shaded regions are excluded.



Collider		$ \lambda_{ee} $	$ \lambda_{e\mu} $	$ \lambda_{\mu\mu} $
LHC	syst. error 0%	1.35	0.95	1.07
	syst. error 10%	1.38	1.00	1.13
HL-LHC	syst. error 0%	0.68	0.51	0.57
	syst. error 10%	0.76	0.68	0.70

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- NSI's obvious targets to scrutinize.
- Example 1: a UV complete model Z'
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 - $M_{Z'} \sim 20 \text{ MeV} - 1 \text{ GeV}$: COHERENT Lar
 - $M_{Z'} \sim 1 \text{ GeV} - 20 \text{ GeV}$: LHCb, CMS 4-leptons
 - $M_{Z'} \sim 20 \text{ GeV} - 4 \text{ TeV}$: HL-LHC 2-leptons
 - There are parameter regions for correlated studies

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Low-energy expts \leftrightarrow LHC complementary!

